

# Bidimensional integration for powder diffraction

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# Outline of the talk

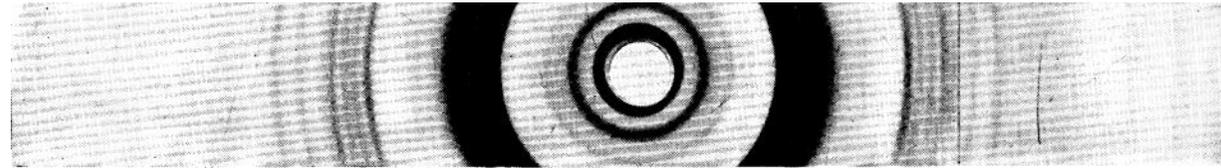
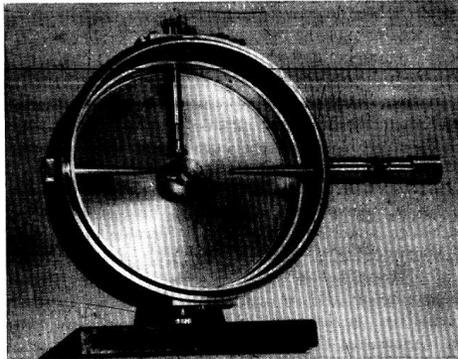
- Principle and comparison for bidimensional detector diffraction
- Main step during integration
- Hand on Tutorial with Fit2D

# Preamble

- The use of two-dimensional detectors with synchrotron radiation enjoys a growing popularity.
- It is now possible to record the entire Debye - Scherrer rings at high angle and good resolution from few minutes to fractions of second.
- The field of applications is vast including texture analysis (Wenk & Grigull, 2003) and in situ powder diffraction measurements in dependence on pressure (Hanfland et al., 1999), temperature (Norby, 1997), chemical composition (Meneghini et al., 2001),

# Chronology

First bidimensional detector for powder diffraction film is the film in Debye-Scherrer camera.



First experiments beginning of the nineties, only utilized thin equatorial strips (Norby, 1997).

Introduction of freely available software the integration of the entire image to a standard one-dimensional powder diffraction pattern (Hammersley et al., 1996)

# Two-dimensional detector

## Cylindrical

- ☺ High angular range
- ☹ Fixed and ☺ constant spatial resolution
- ☹ Complicated and fixed experiment geometry

Maud could handle it!!!

## Flat

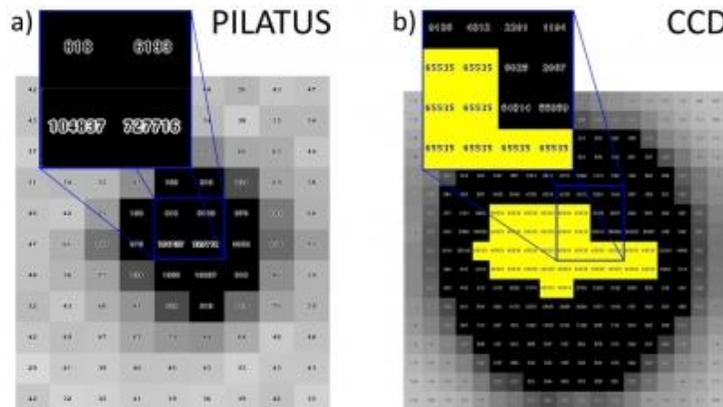
- ☹ Size dependent angular range (limited)
- ☺ Variable spatial resolution distance sample detector
- ☺ Easy experiment geometry
- ☺ Cheaper (IP 70-100 k€)

# Calibration & Integration

Extraction of standard powder diffractograms from two-dimensional images requires knowledge of the diffraction and azimuthal angle at each pixel.

# Detector characteristics

- **The dynamic range** of the detector limits the intensity differences that are recordable on one image. The higher the dynamic range stronger is the available contrast. (bigger grey scale for your image)

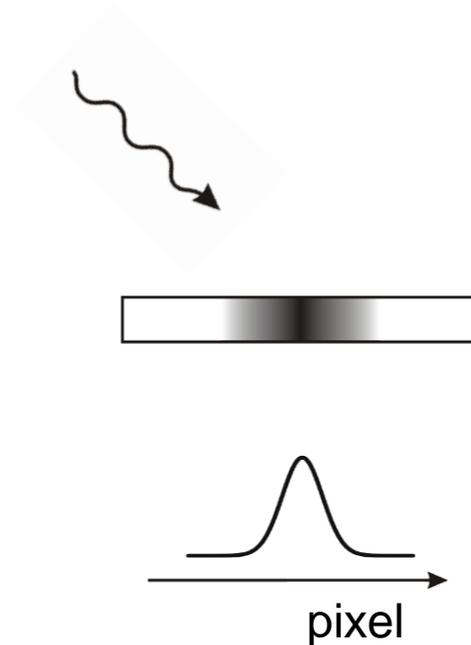
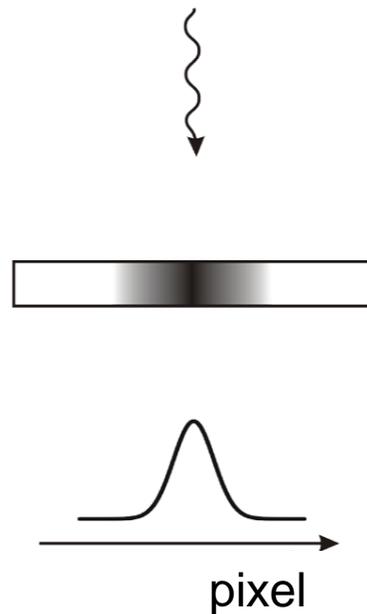


- **The detective quantum efficiency (DQE)** is a measure of the detector's ability to reduce the noise ratio related to detector effectiveness.

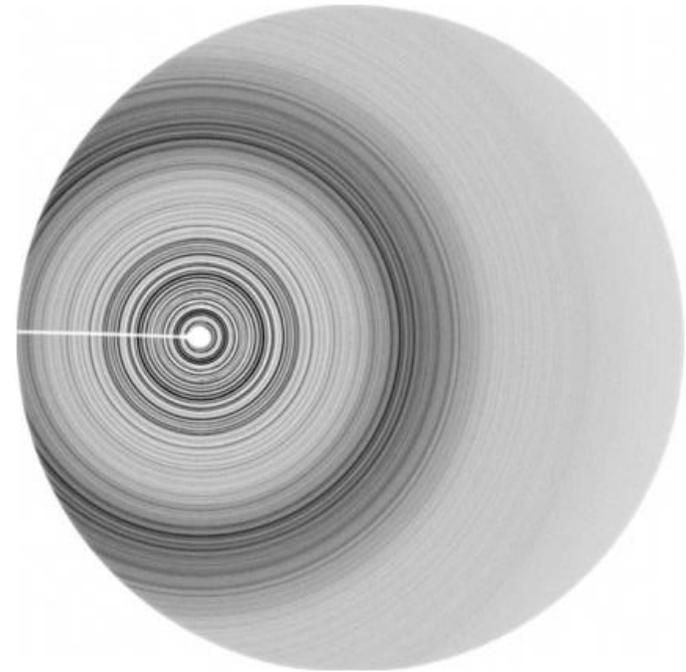
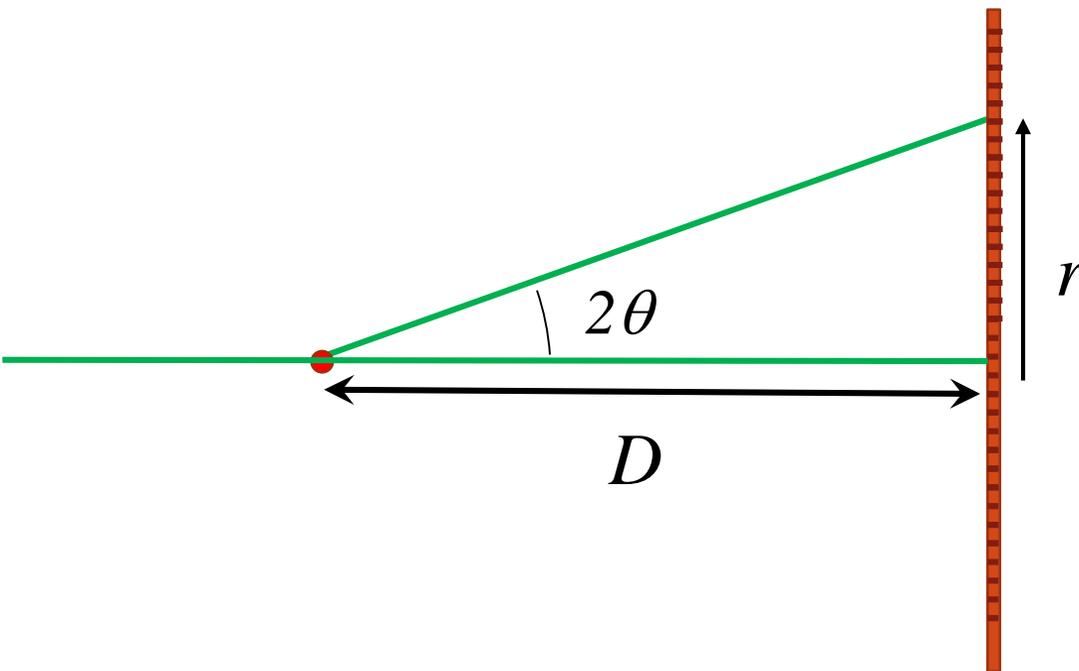
$$DQE = \frac{\left[ \frac{I_{out}^2}{\sigma_{out}^2} \right]}{\left[ \frac{I_{in}^2}{\sigma_{in}^2} \right]}$$

# Detector characteristics

- **Point spread function (PSF)** – the detectors signal to a delta function stimulus. Generally give the signal a Gaussian spread.

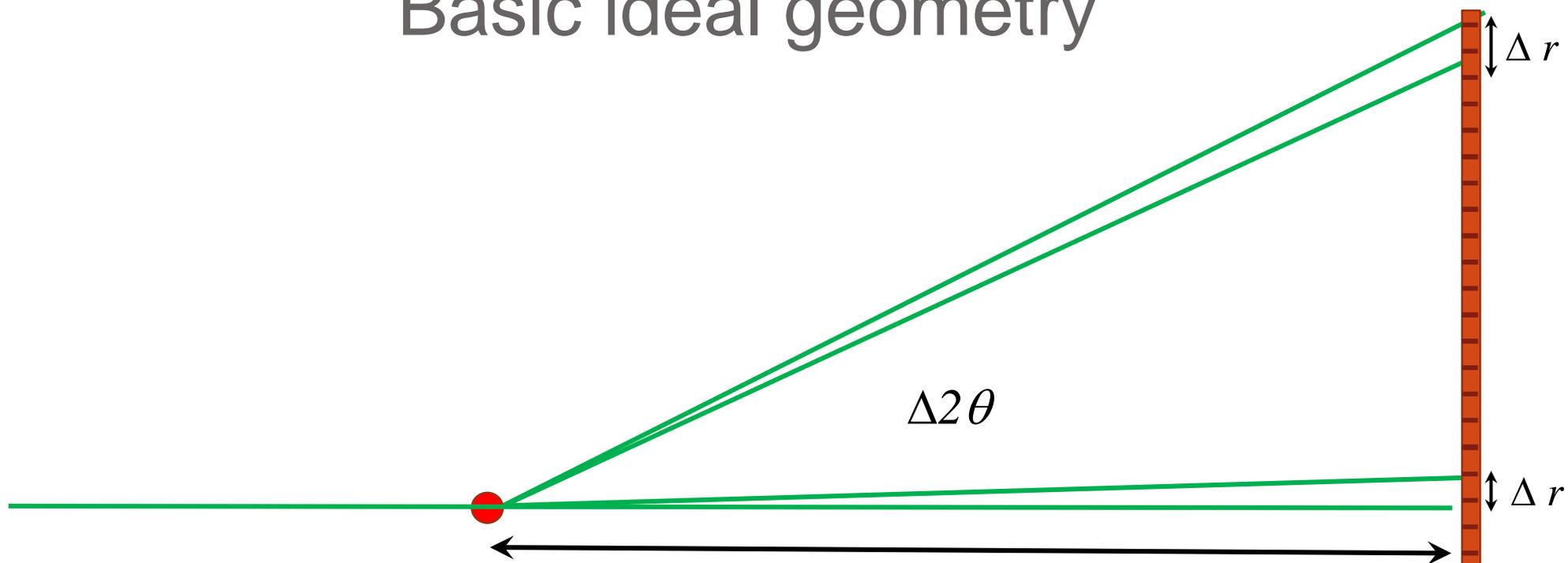


# Basic ideal geometry



$$D \tan 2\theta = r$$

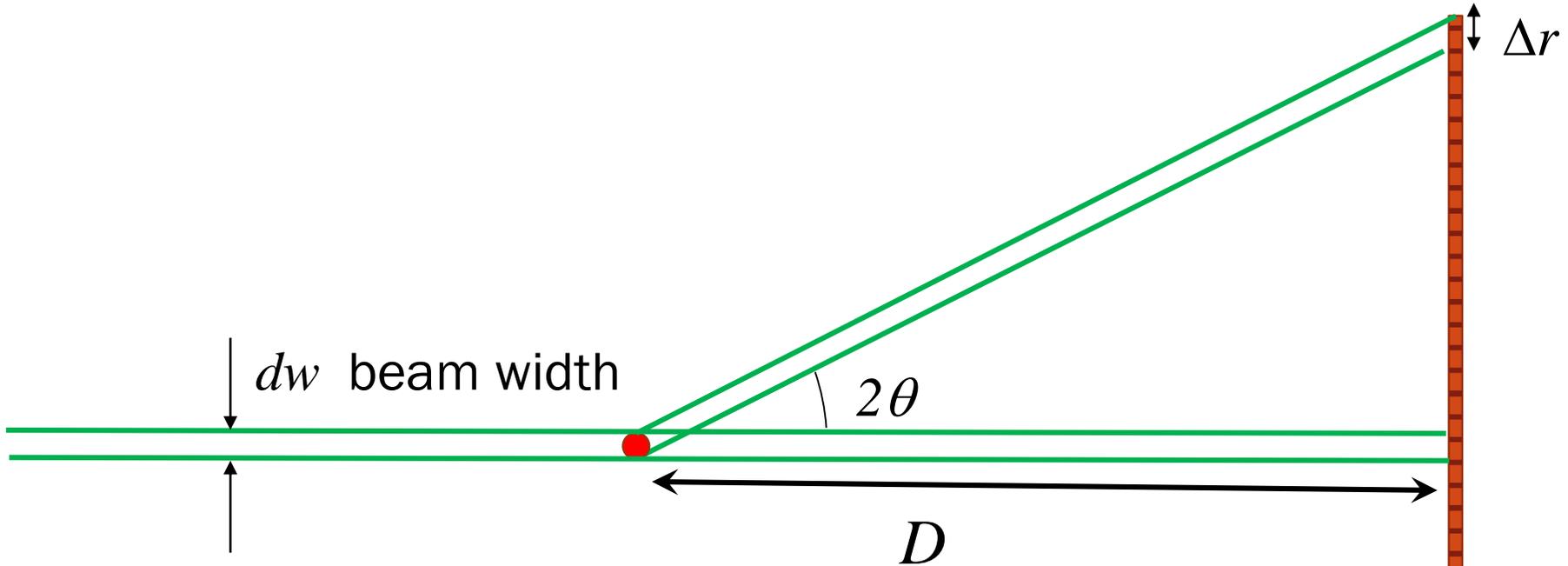
# Basic ideal geometry



$$D \tan 2\theta = r$$

$$\frac{D}{\cos^2(2\theta)} d2\theta = dr$$

# Basic ideal geometry finite beam



$$\frac{dw}{\cos(2\theta)} = dr$$

# Basic ideal geometry finite beam

$$\frac{dw}{\cos(2\theta)} = dr$$

larger footprint on the image plate at higher  $2\theta$

‘number of points (pixels) across the peak... is not changed’ (Norby, 1997).

When using focusing optics (von Dreele et al., 2006) the detector distance to the optics is fixed and the focal spot of the beam well below the PSF of the detector. The resolution is governed solely by the PSF of the detector.

PSF IP  $\sim 300 \mu\text{m}$

# Basic ideal geometry finite beam

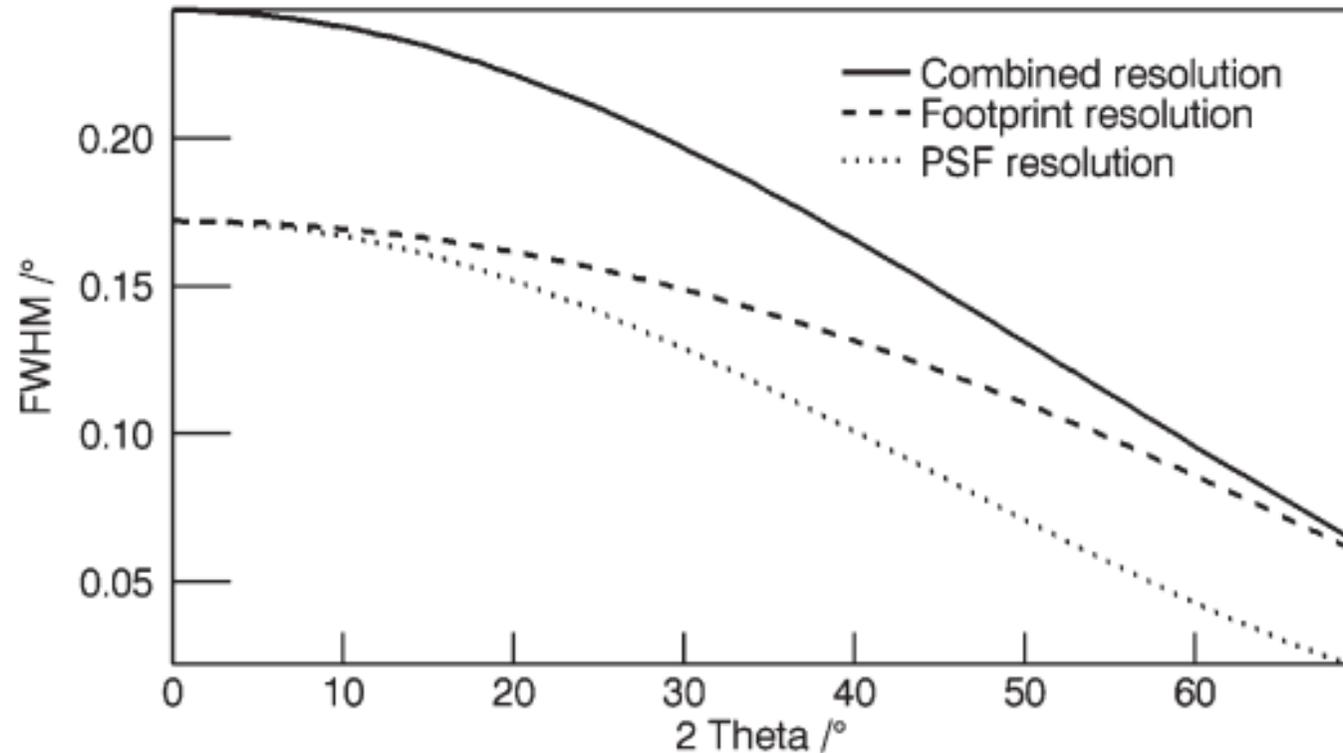
$$D \tan(2\theta + \text{FWHM}) = r + \Delta r$$

$$D \tan(2\theta + \text{FWHM}) = D \tan(2\theta) + \Delta r$$

$$\text{FWHM} = \arctan\left(\tan(2\theta) + \frac{\Delta r}{D}\right) - 2\theta$$

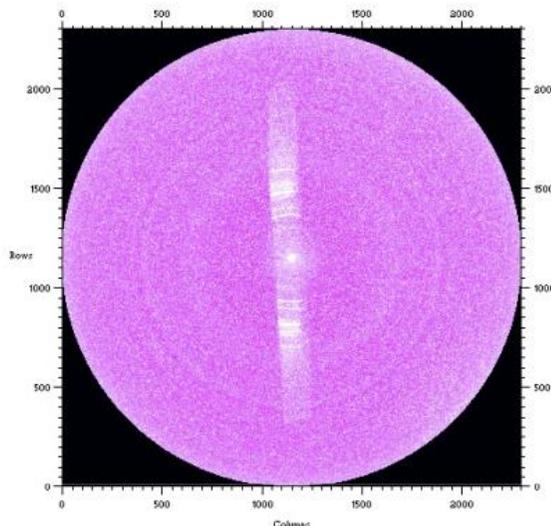
$$\Delta r = \text{PSF} * \frac{dw}{\cos(2\theta)} \quad \text{If both Gaussian} \quad \Delta r = \sqrt{\text{PSF}^2 + \frac{dw^2}{\cos^2(2\theta)}}$$

# FWHM

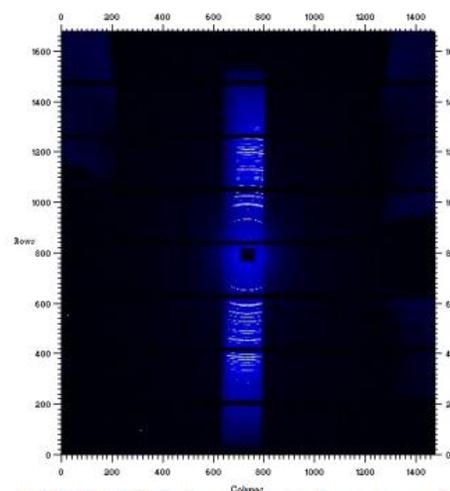


An example of the resolution contributions of an ideally aligned typical image plate detector placed 100 mm from the sample. The footprint of a 0.3 mm diffracted beam and the PSF of 0.3 mm of the detector contribute varying amounts to the resolution over the diffraction angle range.

# Area Detector comparison



Mar345 IP detector



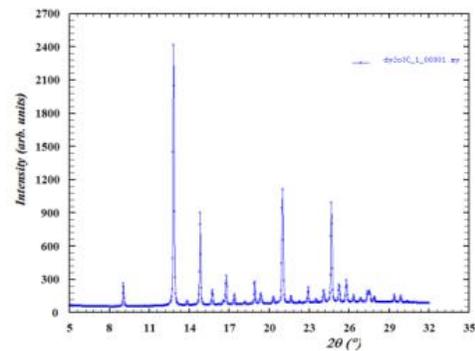
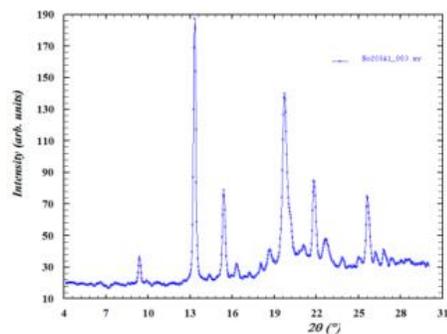
CMOS hybrid-pixel  
technology Pilatus  
2M detector

Pilatus

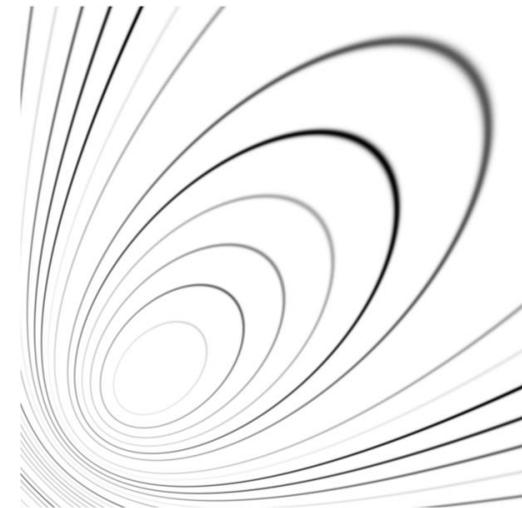
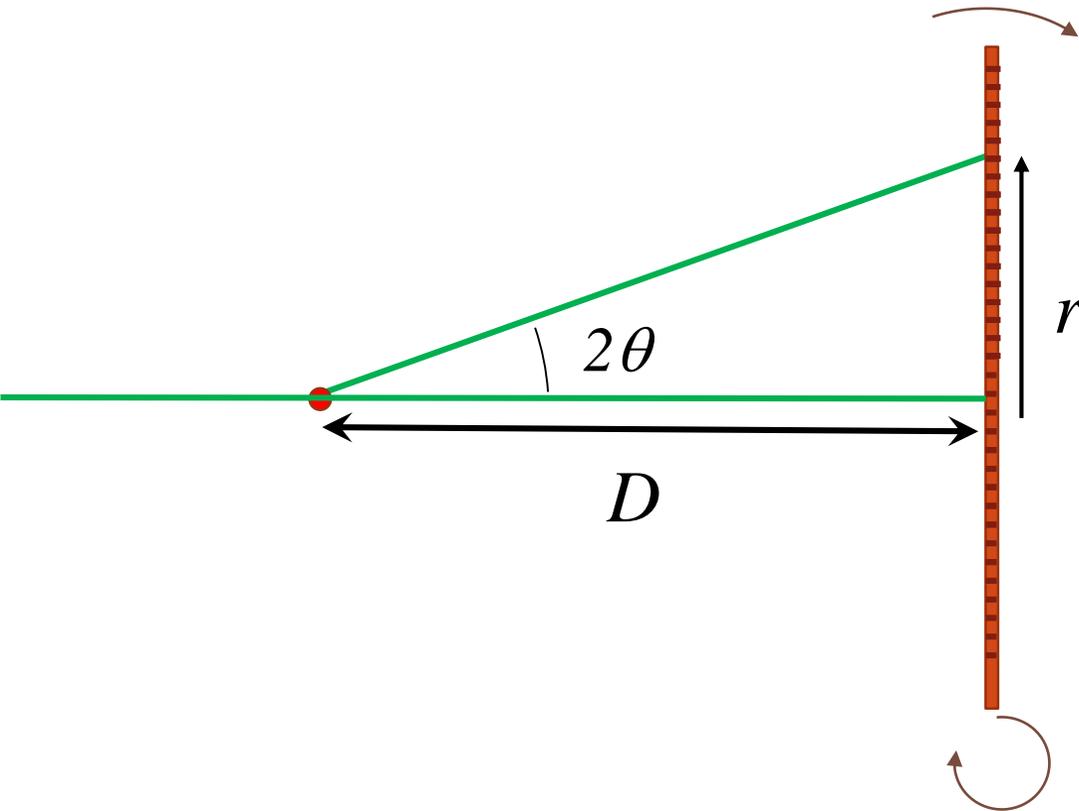
pixel 172  $\mu\text{m}$   
PSF 1 pixel

Mar345

pixel 172  $\mu\text{m}$   
PSF 300  $\mu\text{m}$

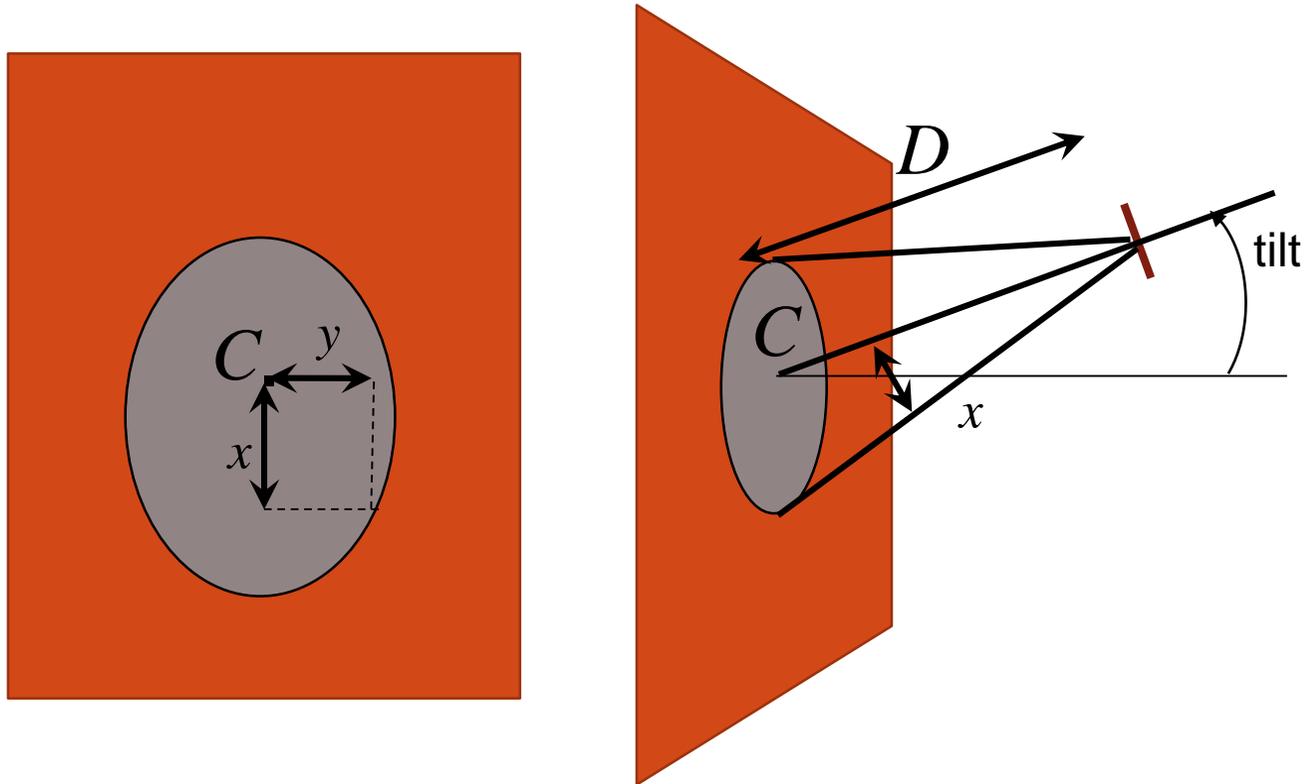


# Tilted detector



© powder3D\_ip manual

# Geometrical convention



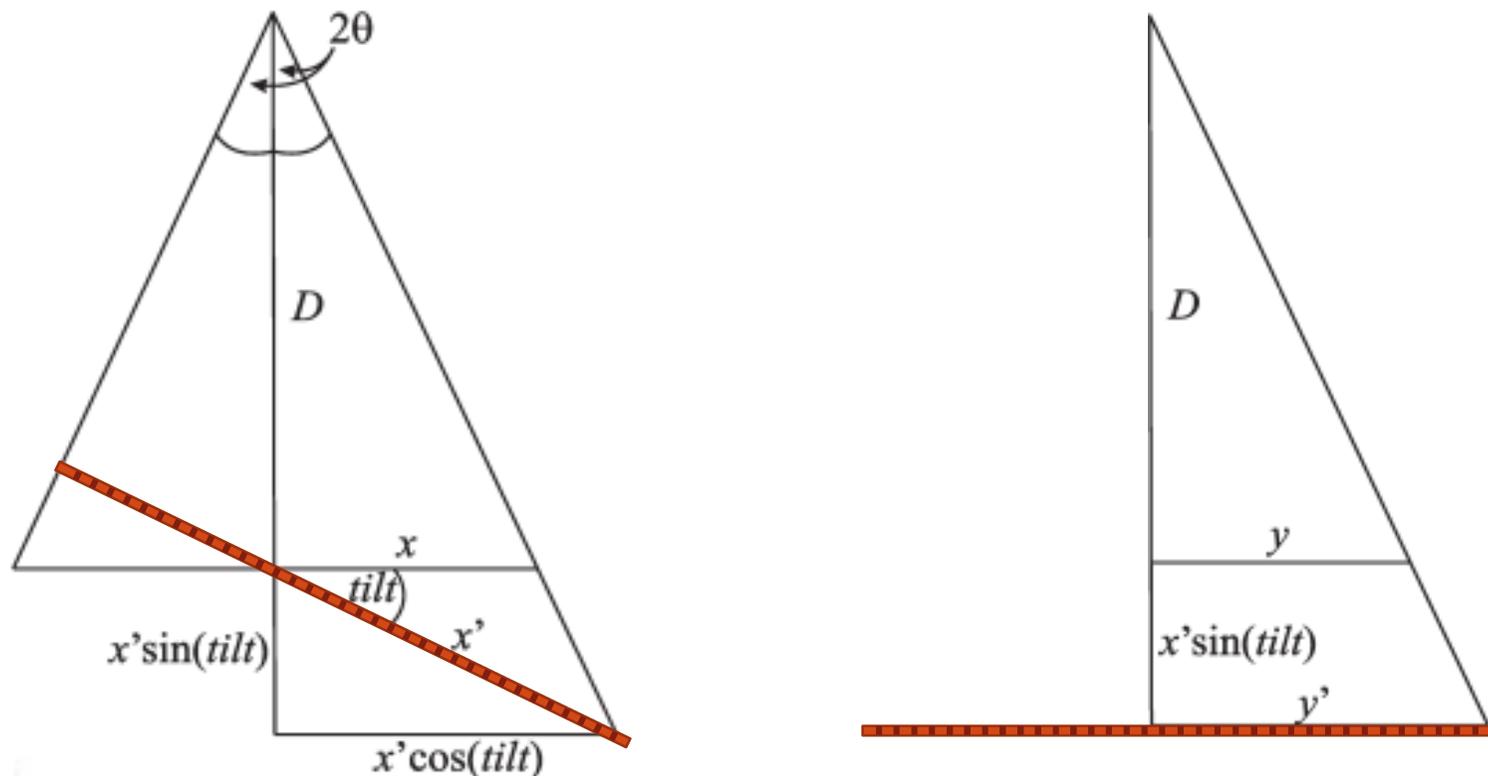
If the cone coordinate system are unaffected by detector tilt previous equation still keep when  $D$  is the distance between sample and the perpendicular plane respect cone axis in which points lay

$$D \tan 2\theta = r$$



$$x^2 + y^2 = D^2 \tan^2 2\theta$$

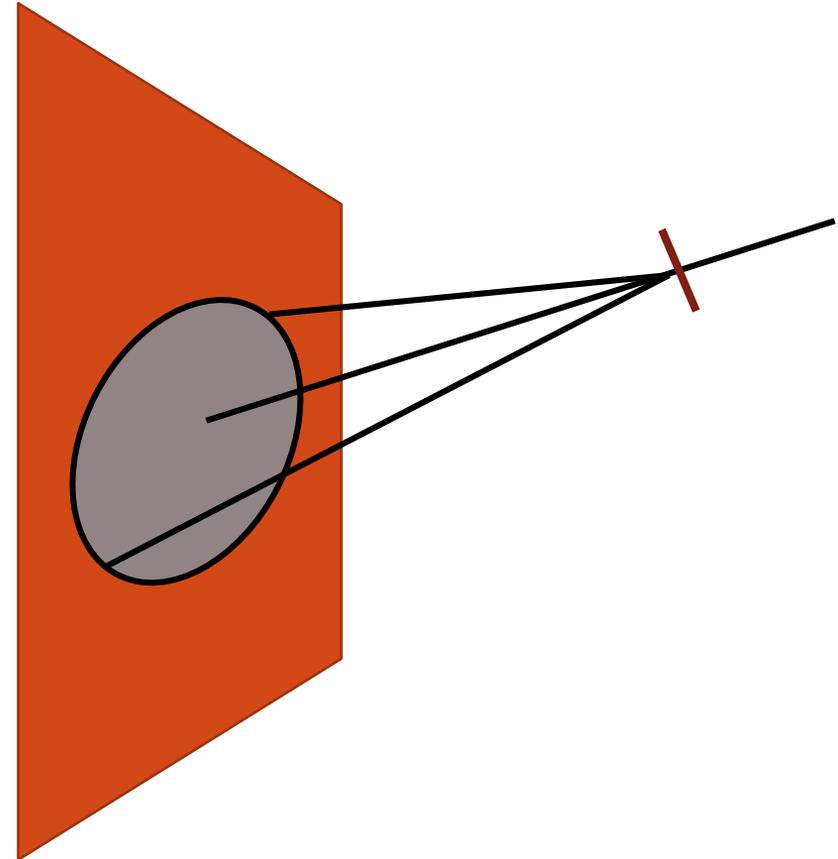
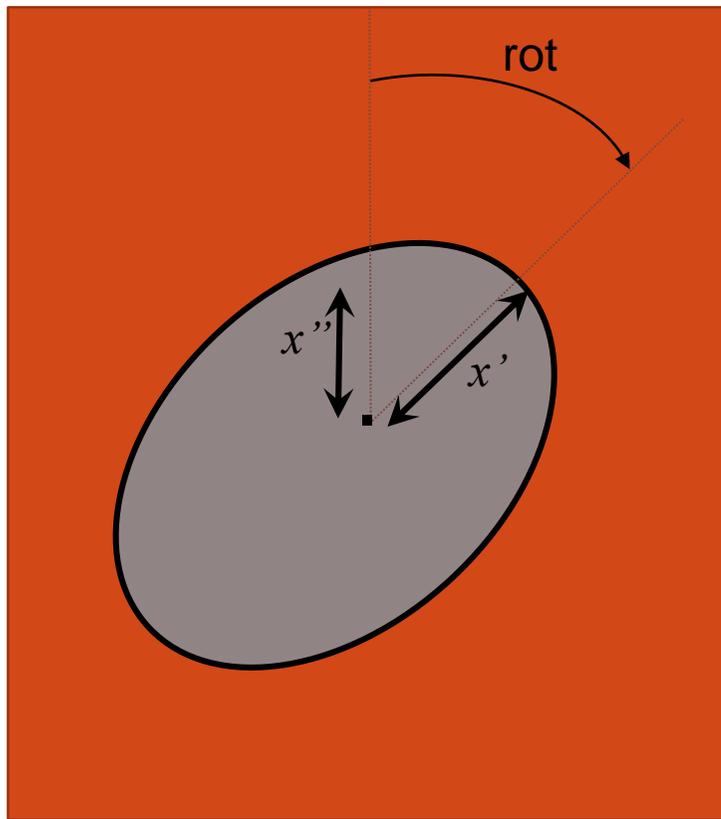
# Diraction Angle Transformation



$$x^2 + y^2 = \left( \frac{Dx' \cos(\text{tilt})}{D + x' \sin(\text{tilt})} \right)^2 + \left( \frac{Dy'}{D + x' \sin(\text{tilt})} \right)^2 = D^2 \tan^2 2\theta$$

$$x'^2 \cos^2(\text{tilt}) + y'^2 = (D + x' \sin(\text{tilt}))^2 \tan^2 2\theta$$

# Arbitrary tilt



$$\begin{aligned}x' &= x'' \cos(\text{rot}) + y'' \sin(\text{rot}) \\y' &= y'' \cos(\text{rot}) - x'' \sin(\text{rot})\end{aligned}$$

# Arbitrary tilt

$$x' = x'' \cos(\text{rot}) + y'' \sin(\text{rot})$$

$$y' = y'' \cos(\text{rot}) - x'' \sin(\text{rot})$$

$$x'^2 \cos^2(\text{tilt}) + y'^2 = (D + x' \sin(\text{tilt}))^2 \tan^2 2\theta$$

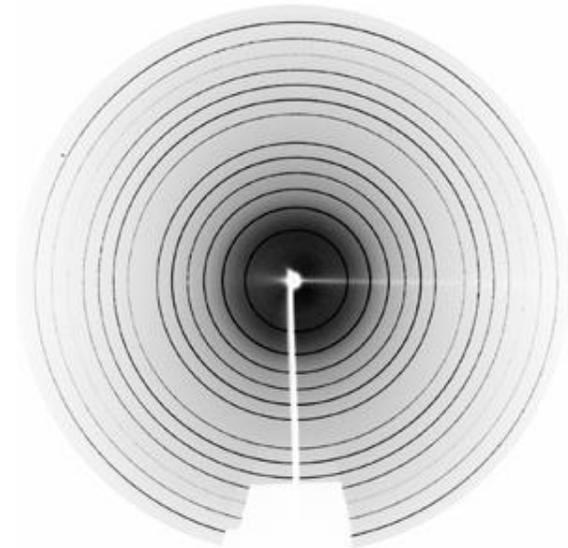
$$\begin{aligned} (x'' \cos(\text{rot}) + y'' \sin(\text{rot}))^2 \cos^2(\text{tilt}) + (y'' \cos(\text{rot}) - x'' \sin(\text{rot}))^2 \\ = (D + (x'' \cos(\text{rot}) + y'' \sin(\text{rot})) \sin(\text{tilt}))^2 \tan^2 2\theta \end{aligned}$$

$$2\theta = \arctan \sqrt{\frac{(x'' \cos(\text{rot}) + y'' \sin(\text{rot}))^2 \cos^2(\text{tilt}) + (y'' \cos(\text{rot}) - x'' \sin(\text{rot}))^2}{(D + (x'' \cos(\text{rot}) + y'' \sin(\text{rot})) \sin(\text{tilt}))^2}}$$

# General Transformations

- We know  $2\theta = f(D, \text{tilt}, \text{rot}, C, x, y)$
- However we don't know *a priori* rot, tilt, D, C

We have an image with some ellipses from a standard



# General Transformations

$a, b = \text{semi-axes}$

$l = \text{semi-latus}$

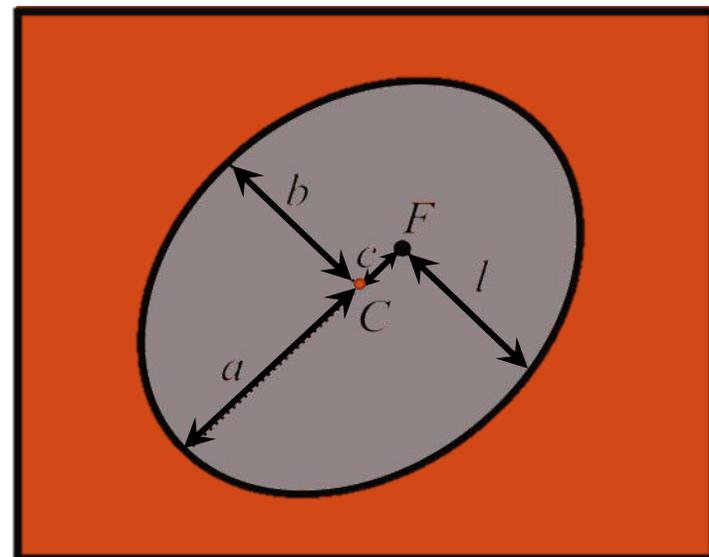
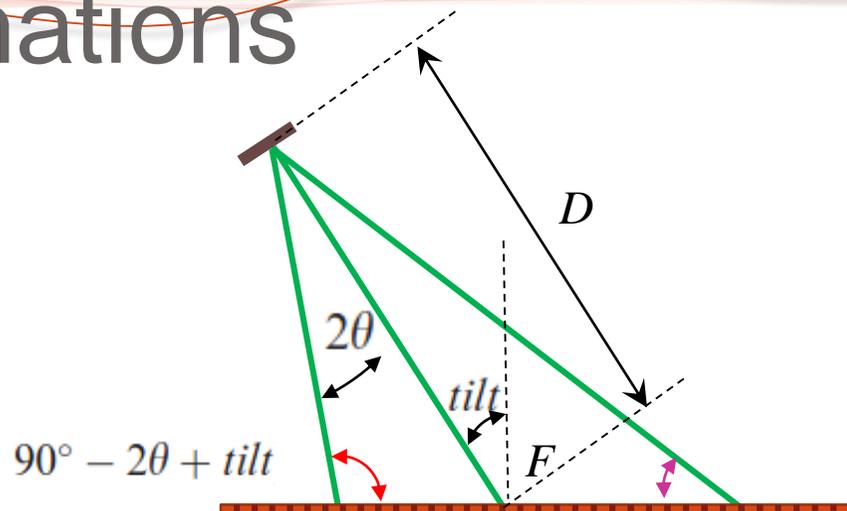
The semi-latus rectum  $l$  is independent of the tilt and only function of  $2\theta$  and detector distance

$$D \tan 2\theta = l$$

Using sine rule

$$\frac{\sin(2\theta)}{a+c} = \frac{\sin(90^\circ - 2\theta - \text{tilt})}{D} \quad c = -a + D \sec(\text{tilt} + 2\theta) \sin 2\theta$$

$$\frac{\sin(2\theta)}{a-c} = \frac{\sin(90^\circ - 2\theta + \text{tilt})}{D} \quad c = a - D \sec(\text{tilt} - 2\theta) \sin 2\theta$$



# General Transformations

$a, b = \text{semi-axes}$   
 $e = \text{eccentricity}$   
 $L = \text{semi-latum}$

$$\frac{D \cos(\text{tilt}) \sin 4\theta}{\cos(2\text{tilt}) + \cos 4\theta} = a$$

$$\frac{2D \sin(\text{tilt}) \sin^2 2\theta}{\cos(2\text{tilt}) + \cos 4\theta} = c$$

$$D \tan 2\theta = l$$

Ellipses identity

$$al = b^2$$

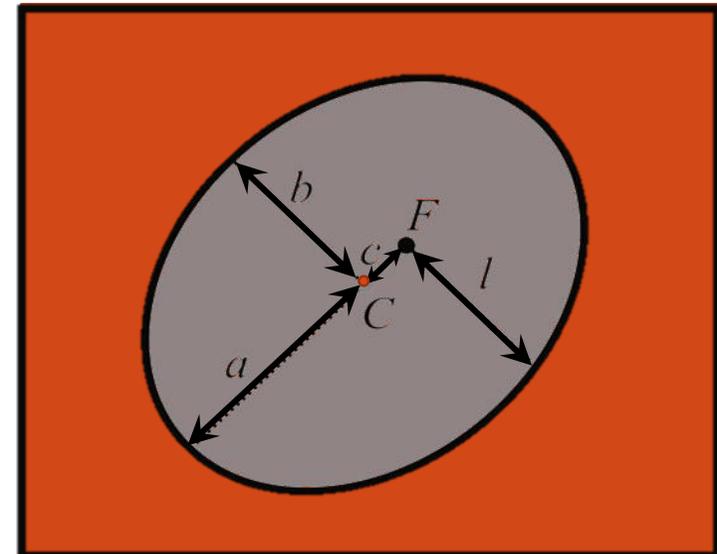
$$e = c/a$$

$$e = \sqrt{1 - \frac{b^2}{a^2}}$$

$$c = \sqrt{a^2 - b^2}$$

$$\frac{D \sqrt{\cos(\text{tilt})} \sqrt{\sin 2\theta} \sqrt{\tan 2\theta}}{\sqrt{\cos(2\text{tilt}) + \cos 4\theta}} = b$$

$$\tan(\text{tilt}) \tan 2\theta = e$$



# Detector Coordinate Transformations

## Ellipses identity

$$\begin{aligned}
 a &= b^2 \\
 e &= c/a \\
 e &= \sqrt{1 - \frac{b^2}{a^2}} \\
 c &= \sqrt{a^2 - b^2}
 \end{aligned}$$

$$\frac{D \cos(\text{tilt}) \sin 4\theta}{\cos(2\text{tilt}) + \cos 4\theta} = a$$

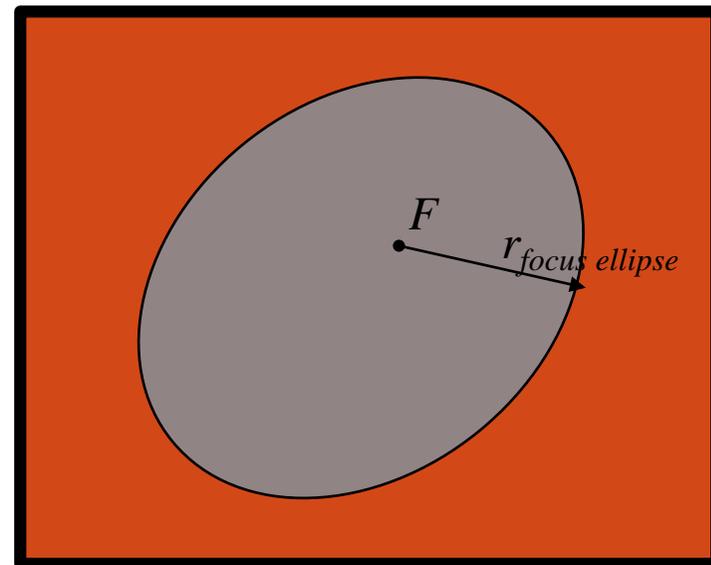
$$\frac{D \sqrt{\cos(\text{tilt})} \sqrt{\sin 2\theta} \sqrt{\tan 2\theta}}{\sqrt{\cos(2\text{tilt}) + \cos 4\theta}} = b$$

$$\tan(\text{tilt}) \tan 2\theta = e$$

$$r_{\text{focus ellipse}} = \frac{a(1 - e^2)}{1 + e \cos(\alpha)} = \frac{D \tan 2\theta}{1 + \cos \alpha \tan(\text{tilt}) \tan 2\theta}$$

$$y_d = r_{\text{focus ellipse}} \sin(\alpha - \text{rot})$$

$$x_d = r_{\text{focus ellipse}} \cos(\alpha - \text{rot})$$

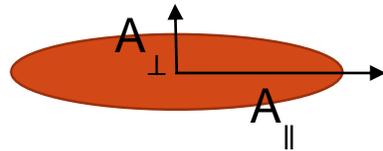


# Intensity corrections

- Polarization correction
- Geometrical correction
- Lorentz correction

# Polarization correction

- The polarization factor arises from partial polarization of the electromagnetic wave after scattering.
- Polarized beam can be represented by two components:  $A_{\parallel}$  and  $A_{\perp}$



- The diffracted intensity is proportional to the square projections of  $\epsilon A$  on  $\epsilon k_{\text{out}}$ , ( $A^2$  for  $\epsilon A \perp \Delta k$ ,  $\cos^2(2\theta)A^2$ )
- The overall intensity reduction of a not polarized beam scattered

$$P \propto \frac{1 + \cos^2 2\theta}{2}$$

# Polarization correction

the primary beam be polarized this changes the correction to

$$P = P_0 - P'$$

$$P' = \frac{1}{2} J' \cos 2\alpha \sin^2 2\theta$$

Where  $\alpha$  is the azimuthal angle,  $\alpha=0$  equatorial  $J'$  is the degree of polarization

$$J' = \frac{(A_{\parallel} - A_{\perp})}{(A_{\parallel} + A_{\perp})}$$

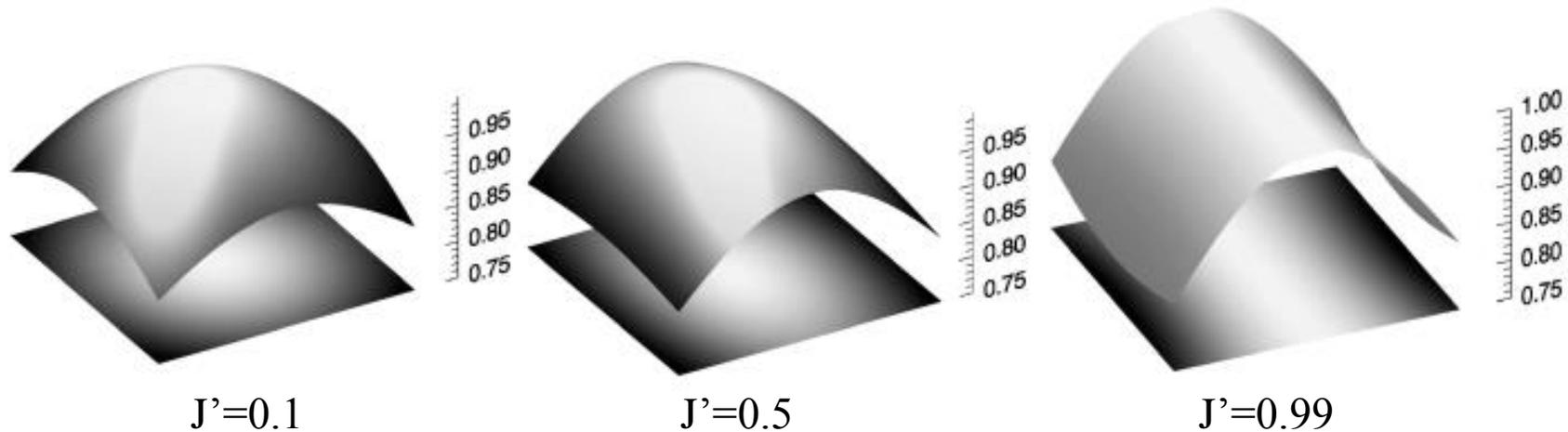
$J'=1$   $\parallel$  polarization

$J'=-1$   $\perp$  polarization

$J'=0$  no polarization

# Polarization correction

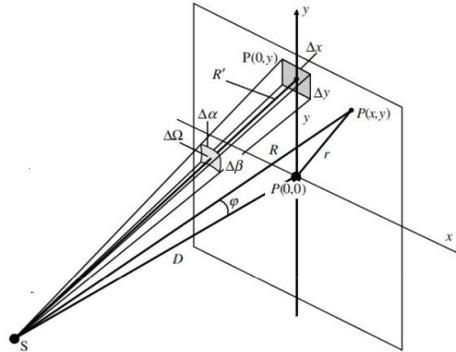
$$P = P_0 - P'$$



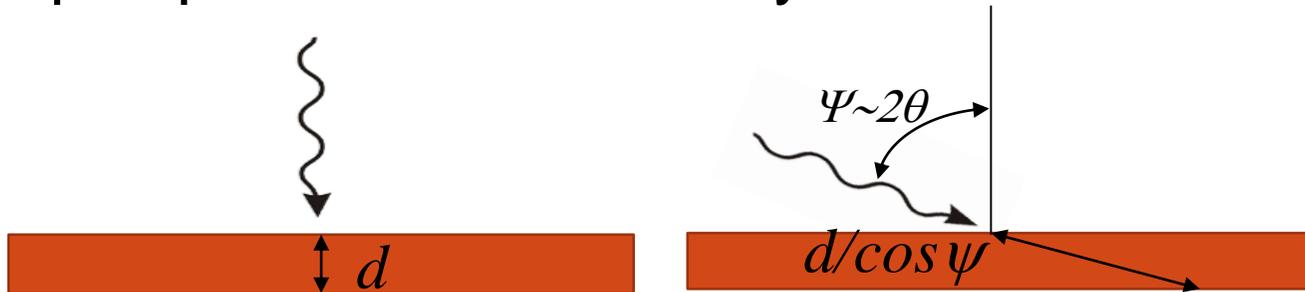
Two dimensional polarization correction for an ideally aligned detector. The correction is applied by dividing the intensities by the displayed correction values.

# Geometrical correction

- Change of solid angle cover by a pixel

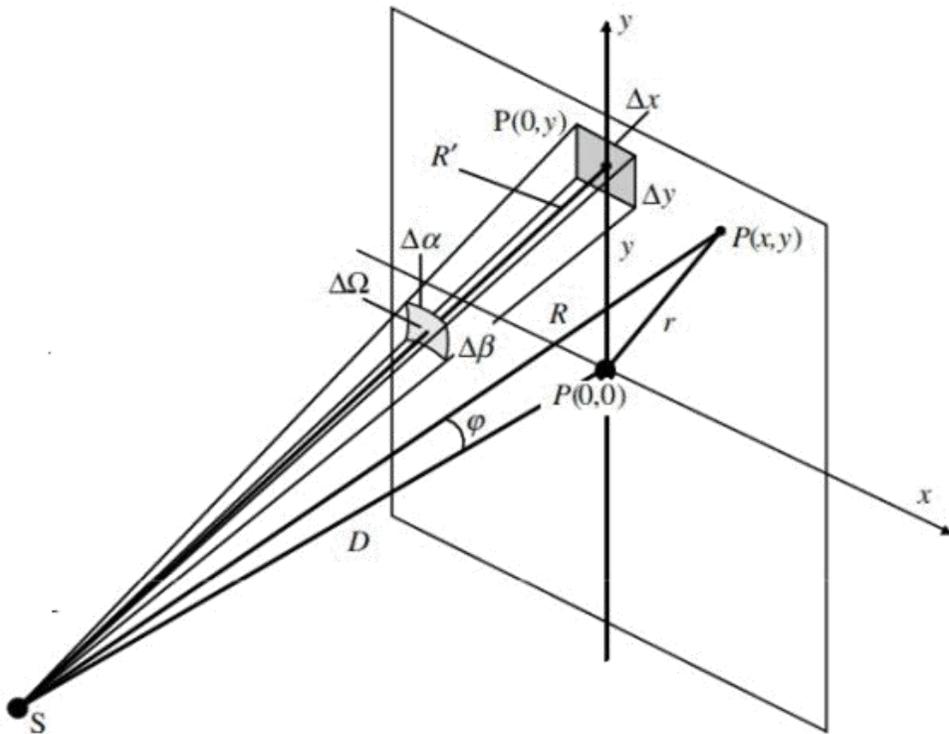


- partial transmission of X-ray through the active layer  
Optic path in the active layer



# Geometrical correction

- Change of solid angle cover by a pixel



$$\Delta\beta = \frac{D}{D^2 + y^2} \Delta y = \frac{D}{R'^2} \Delta y \quad \Delta\alpha = \frac{\Delta x}{R'}$$

$$\Delta\Omega = \Delta\beta\Delta\alpha = \frac{D}{R'^3} \Delta A$$

*Flux for pixel*  $\propto \Delta\Omega$

$$K = \frac{\text{Flux}(x,y)}{\text{Flux}(0,0)} = \frac{D^3}{R'^3} = \cos^3\psi \cong \cos^3 2\theta$$

Implemented on fit2D  
geometrical correction to intensity

# Geometrical correction

- K is the angular dependent detector efficiency

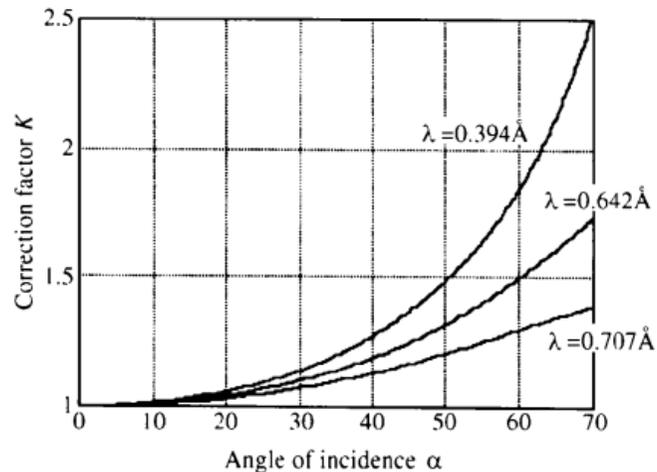


Fig. 2. Correction factor  $K$  for three different wavelengths, such  $I_{corr} = I_{obs}/K$ .

$$K = \frac{1 - e^{-\frac{\mu d}{\cos 2\theta}}}{1 - e^{-\mu d}} = \frac{1 - e^{-\frac{\ln T_{\perp}}{\cos 2\theta}}}{1 - T_{\perp}}$$

$$I_{corr} = I_{obs} / K$$

Implemented in Fit2D inside flat field correction

# Lorentz correction

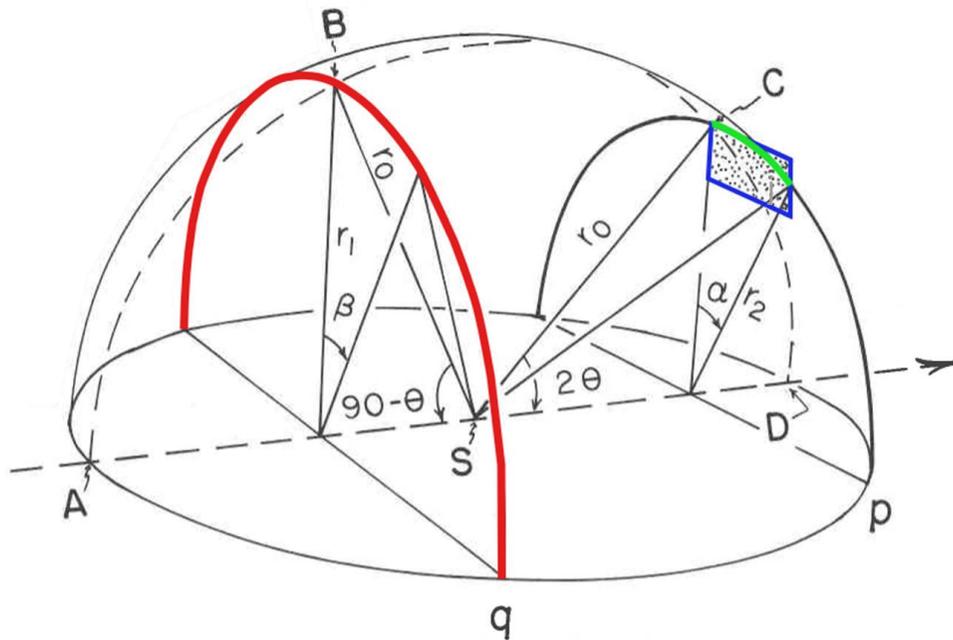
For Debye-Scherrer and Bragg-Brentano classical term implemented in Rietveld software (ignoring all constants)

$$L = \frac{1}{\sin\theta \sin 2\theta} \quad \text{or} \quad \frac{1}{\sin^2\theta \cos\theta}$$

This factor include 3 effect

- **Statistic**
- **Slits acceptance**
- **Angular velocity**

# Lorentz correction statistic and slits factor



Crystallite in diffraction condition  
 $\propto B \text{ arc} \propto \sin 90-\theta = \cos \theta$

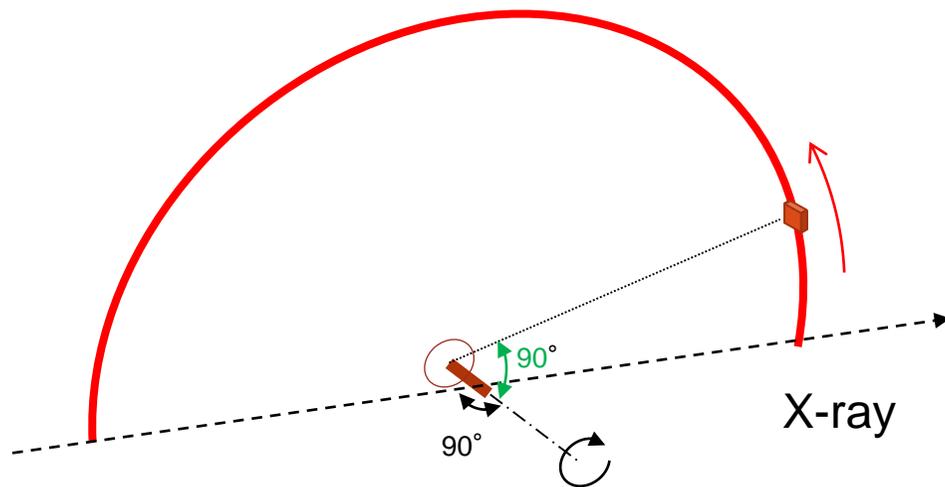
Slits acceptance  
 $\propto \alpha = \arcsin(j/r_0 \sin 2\theta) \approx j/r_0 \sin 2\theta$

$$Lp \propto \cos \theta / \sin 2\theta = 1/2 \sin \theta$$

# Single crystal angular velocity

For Debye-Scherrer (equatorial scan) Ewald sphere velocity reduces to Darwin form

$$L \approx 1/\sin 2\theta$$



But for two dimensional detector the angle between sample axis rotation and detector plane is different from 90°

# Non equatorial Lorentz correction

$$L \sim \frac{1}{\cos \mu \cos \nu \sin T}$$

$\mu$  angle between axis of sample rotation and plane  $\perp$  beam  
 $\gamma$  and  $\nu$  horizontal and vertical reflection displacement

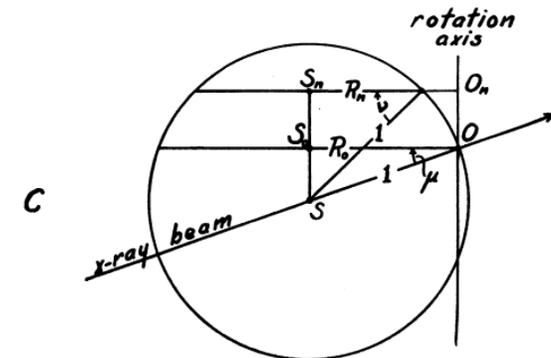
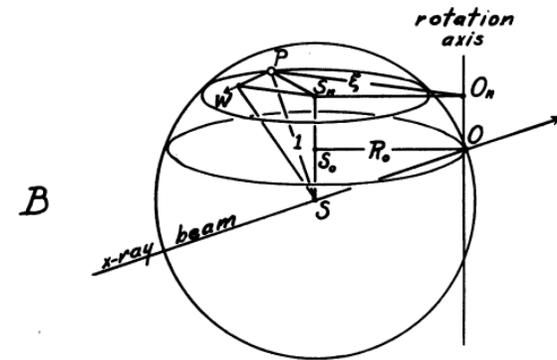
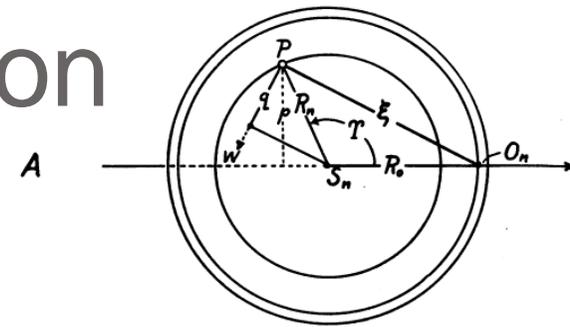


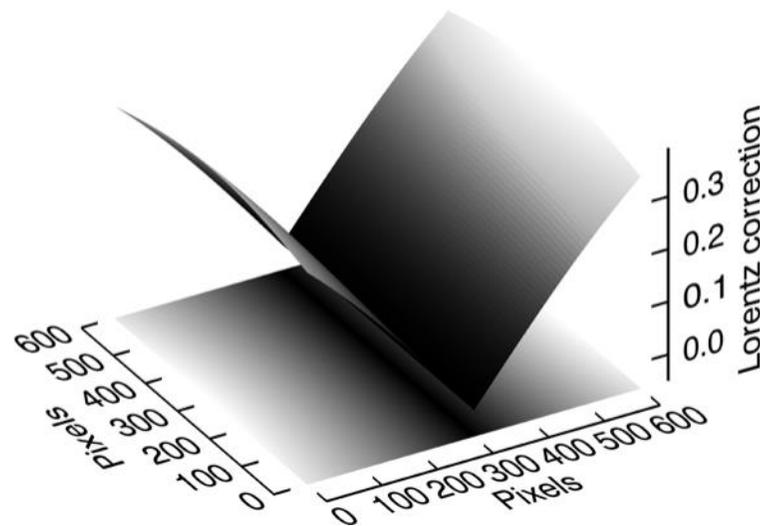
FIGURE 1

# Non equatorial Lorentz correction

Expressed in terms of azimuthal angle

$$L^{-1} = \cos \mu \frac{\cos \alpha \tan 2\theta}{\sqrt{1 + \cos^2 \alpha \tan^2 2\theta} \sqrt{1 + \sin^2 \alpha \tan^2 2\theta}}$$

Although Fit2D documentation is not clear if apply the azimuthal dependent Lorentz correction is clearly state that intensity are renormalized by a  $\sin 2\theta$  factor to maintain compatibility with current Rietveld software



The two-dimensional single crystal Lorentz correction for an ideally aligned detector. Note the zero values in the central valley. These cause divergent intensities as they are multiplied with the inverse of the Lorentz correction. Therefore the intensities in that region have no meaning. The central valley is parallel to the sample rotation axis.

# Pre treatment of the data

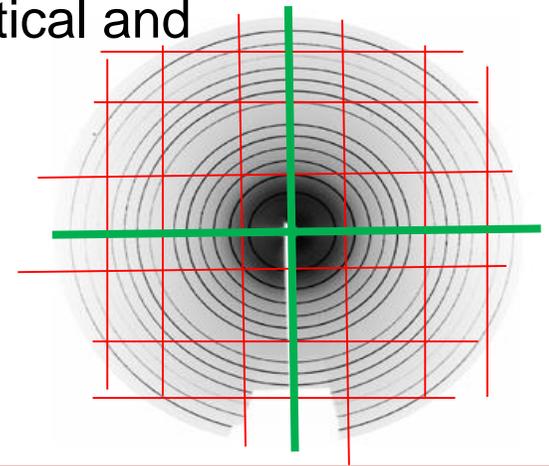
- Several region should be rejected from integration:
  - Beam stop
  - Bad pixel (dust grain, damaged pixel, etc)
  - Dead zone
  - Saturating spot (big crystallite or cluster of oriented crystallite)
- Manually selecting polygons, arcs, circles
  - Fit2D
- Automatic setting threshold on intensity or filtering the distribution

# Calibration

- Definition of a set of d-spacing for calibration image
- Definition of the direct beam position center of circle in orthogonal detector
- Definition of ellipses and experimental geometry
- Refinement
- Azimuthal check

# Center determination

- User assisted method (Fit2D)
  - Direct beam trace (semi opaque beam stop)
  - Circle or ellipse center (click on ellipse)
  - Fit projection
- Fully automatic
  - Cervellino method: minimize the inverse normalized variance of the radial distribution function.
  - Powder3D\_IP Intersection of mirror plane for vertical and horizontal cut (need center to be in the plate)

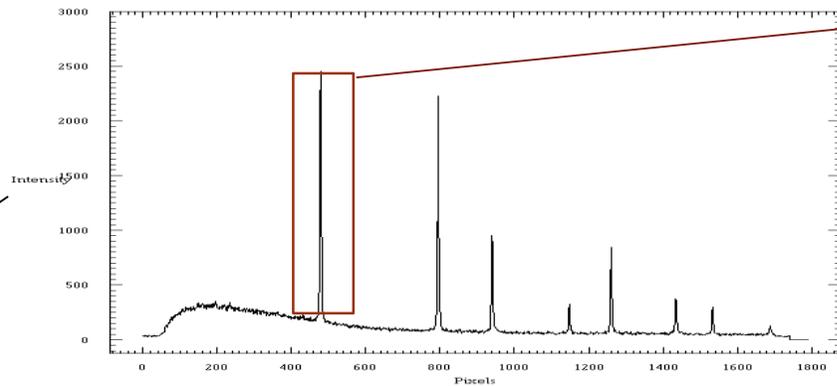
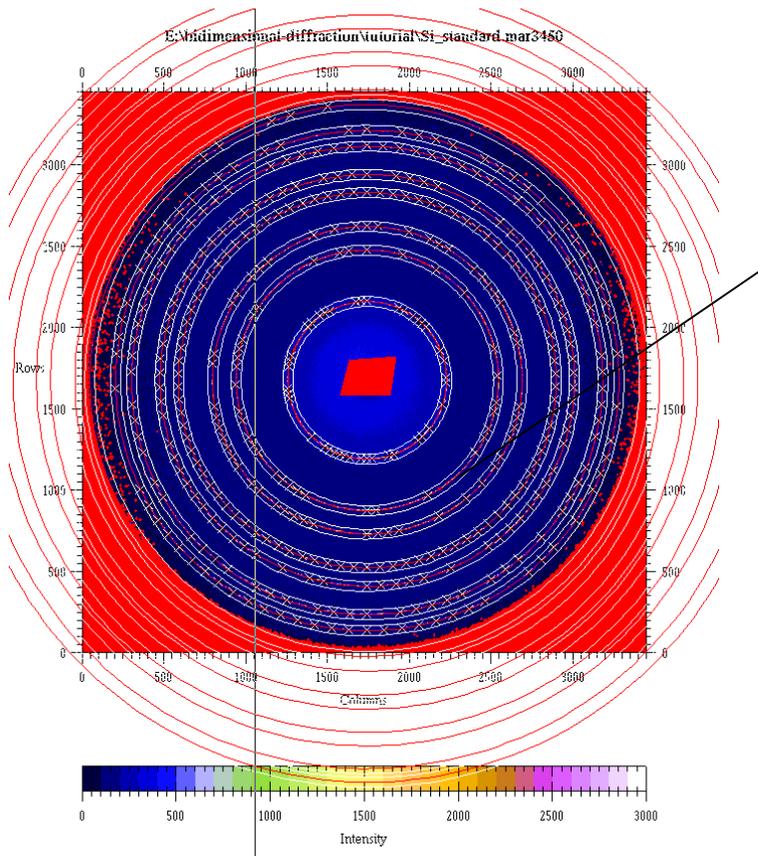


# Tilt and rotation

- Fit2D: manual fit of several ellipses
- Powder 3D: automatic fit with the Hough transform, as has been described by Rajiv and by Fisker

# Refinement

- The radial line intersection method  
Fit2D, Powder3D, Maud



Gaussian fit

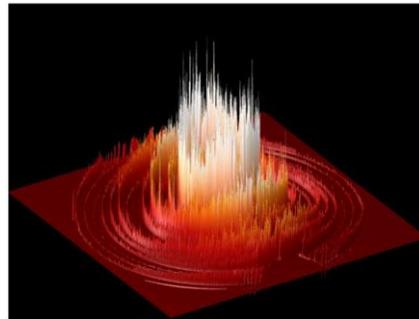
Several slice are cut starting from the approximated center (bins) and successively fitted in terms of center, tilt detector distance etc.

$$r_{focus-ellipse} = \frac{D \tan 2\theta}{1 + \cos \alpha \tan(\text{tilt}) \tan 2\theta}$$

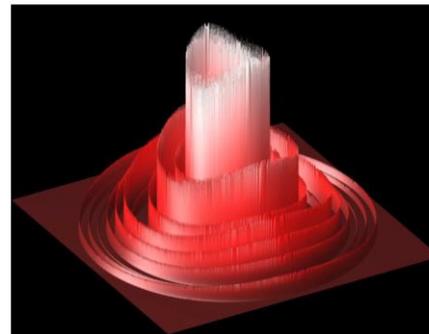
$$w_i = D \sin 2\theta$$

# Whole image refinement

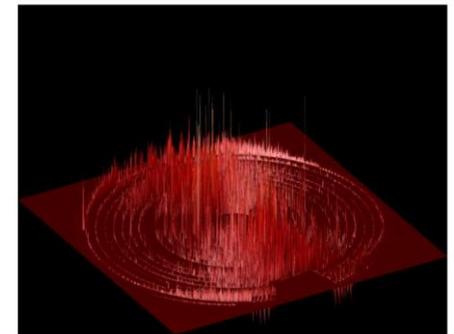
- Powder3D\_IP: the entire image is reconstructed and subtracted from the measured image. This difference is weighted and used as the residual for the refinement process.



exp



fit



residual

# Comparison of methods for no tilt

Software/method	X0[pixels]	Y0[pixels]	Tilt[°]
Real Values	1150	1150	0
Fit2D (traditional)	1150	1150	7.60E-07
Powder3D (traditional)	1150.001	1150.0007	-0.00020402
Powder3D WIR 10*10 binning	1149.9966	1149.9971	-0.000687
Powder3D WIR 2*2 binning	1150.0007	1150.0001	-0.000104
Powder3D WIR no binning	1150.0005	1150	-0.00010347
	Distance[mm]	Wavelength[ Å]	Chi <sup>2</sup>
Real Values	100	1	
Fit2D (traditional)	99.99296	1.000053	4.92E-07
Powder3D (traditional)	100.00384	0.99998171	1.89E-07
Powder3D WIR 10*10 binning	99.999813	0.999989	4.90E-07
Powder3D WIR 2*2 binning	99.999886	0.999999	1.14E-08
Powder3D WIR no binning	99.999881	0.99999842	1.11E-08

# Integration

- In the idea that detector doesn't move: however if you cliche is not too spotty you can still refine tilt and center **Just press a button**
- Two options:
  - Full ring integration (high statistic)
  - Multiple azimuthal bin texture studies
- Take care of consistency between correction done by integration software and rietvend program.

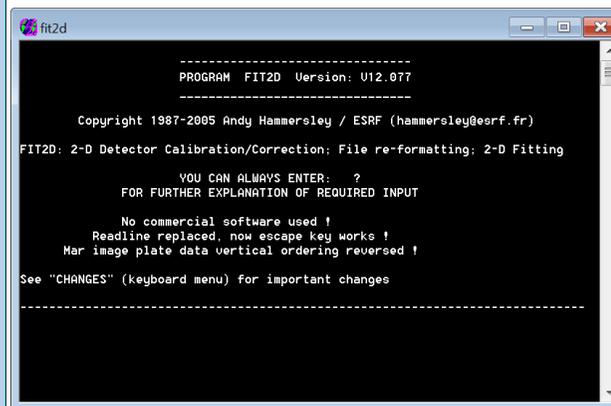
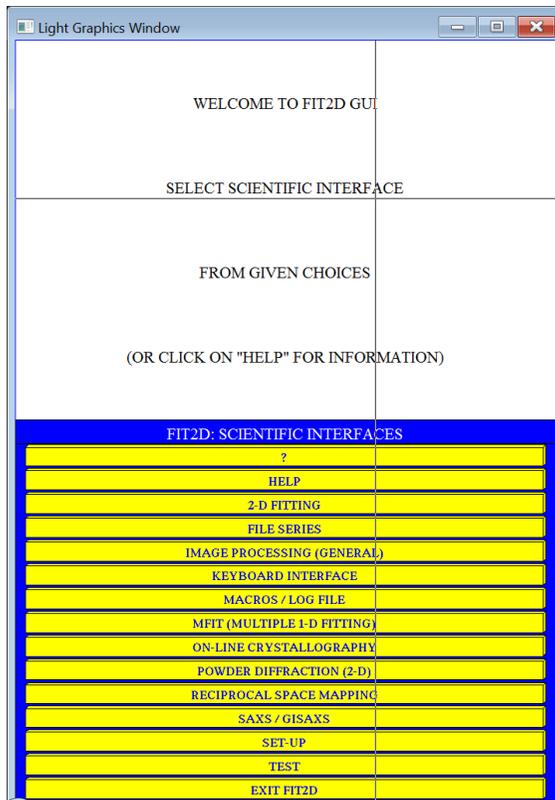
# Final consideration

- Trash in, trash out:
  - if collection on the calibrant is bad
  - calibration is bad and integration on sample is worst

Less is better : don't hesitate to use masking if calibration is reasonable

# Fit2D

- The software is composed of a graphical windows and a shell where information appear and command line mode could activated



- +Very flexible
- +Open a lot of image kinds
- +widely available linux win mac

- obscure documentation
- not frequently update
- limited choice of option

# Start

Light Graphics Window

DIMENSIONS OF PROGRAM ARRAYS  
(need to be big enough to store  
and work on data)

O.K. CANCEL ? HELP INFO

DESCRIPTIONS	VALUES	CHANGE
FIRST DIMENSION OF ARRAYS	5120	X-DIMENSION
SECOND DIMENSION OF ARRAYS	5120	Y-DIMENSION
CREATE MEMORY ARRAYS	YES	MEMORY
CREATE VARIANCE ARRAYS	YES	VARIANCES

Click on variable to change, or 'O.K.'

Define the image array size at list  
the number of pixel in the image

Use of memory array on actual computer  
doesn't matter any limit consideration

Keep track of variance during manipulation

# Start

Light Graphics Window

CONTROL OF DETECTOR

DISTORTION CORRECTIONS

O.K. CANCEL ? HELP INFO

DESCRIPTIONS	VALUES	CHANGE
SUBTRACT DARK CURRENT IMAGE	NO	DARK CURRENT
NAME OF DARK CURRENT FILE	dark_current.bin	DC FILE
APPLY FLAT FIELD CORRECTION	NO	FLAT-FIELD
NAME OF FLAT-FIELD FILE	flat_field.bin	FF FILE
APPLY SCALING AFTER FLAT FIELD CORRECTION	NO	FF SCALE
FLAT FIELD MULTIPLIER TO APPLY	1000.000	FF MULTIPLIER
APPLY SPATIAL DISTORTION CORRECTION	NO	SPATIAL DIS.
NAME OF SPATIAL DISTORTION FILE	spatial.spline	SD FILE

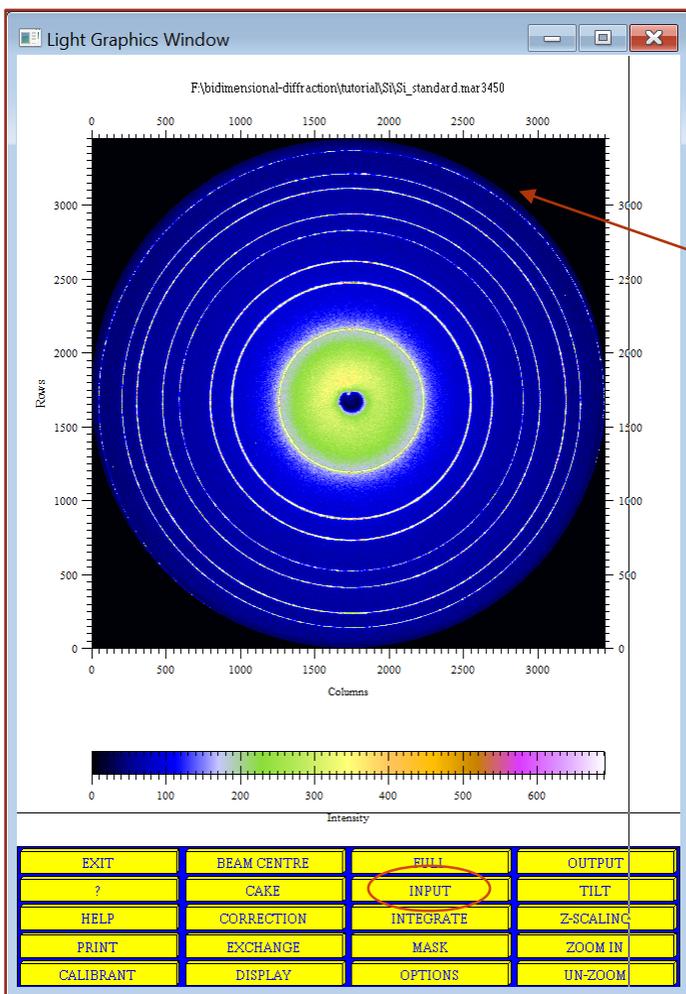
Click on variable to change, or 'O.K.'

- Principal correction
- Dark current image (subtract image)
- Flat field (multiplier image)
- Spatial distortion (spline deforming the image) can include detector transparency

# Start

Just press input and navigate in the menu

Empty region

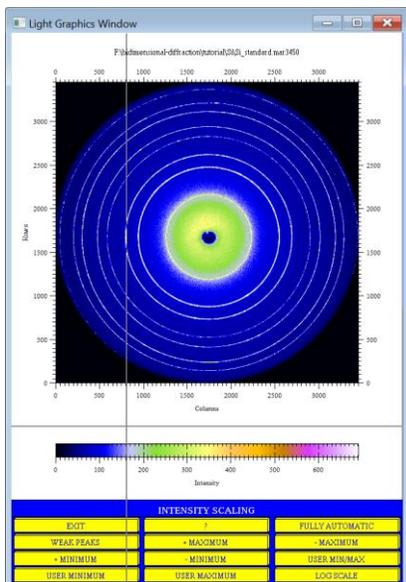


```
fit2d
NOTE: This menu is designed to work on data stored in horizontal
rows. The "POWDER DIFFRACTION" menu NOW PRODUCES 2-THETA,
RADIAL, Q-SPACE DATA IN THE X-DIRECTION. Azimuthal data
is now in the Y-direction. If fitting azimuthal data the
"TRANPOSE" command should be used to convert the data to
row order.

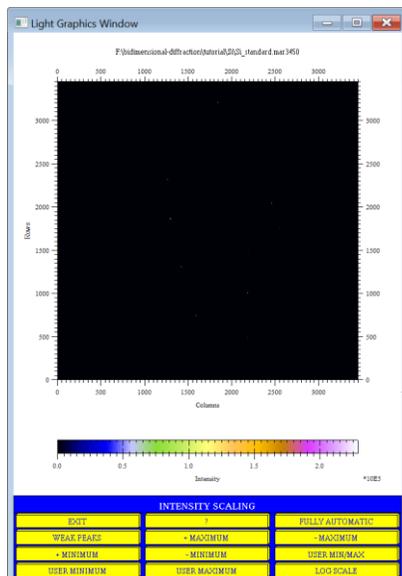
INFO: Image size = 3450 3450 Number of overloaded pixels = 283
INFO: The following information has been extracted from the file header:
Date and time of scan = Thu Apr 25 09:
Sample to detector distance = 36.00 mm
Wavelength is = 0.4100 Angstroms
Phi angle at the start of the oscillation = 0.000 degrees
Phi angle at the end of the oscillation = 3.000 degrees
(These values may be changed using the "GEOMETRY" command.)
INFO: Image size = 3450 3450 Number of overloaded pixels = 283
INFO: The following information has been extracted from the file header:
Date and time of scan = Thu Apr 25 09:
Sample to detector distance = 36.00 mm
Wavelength is = 0.4100 Angstroms
Phi angle at the start of the oscillation = 0.000 degrees
Phi angle at the end of the oscillation = 3.000 degrees
(These values may be changed using the "GEOMETRY" command.)
```

# Z-scaling

- Define color code for intensity fundamental to discover aberration and outlier



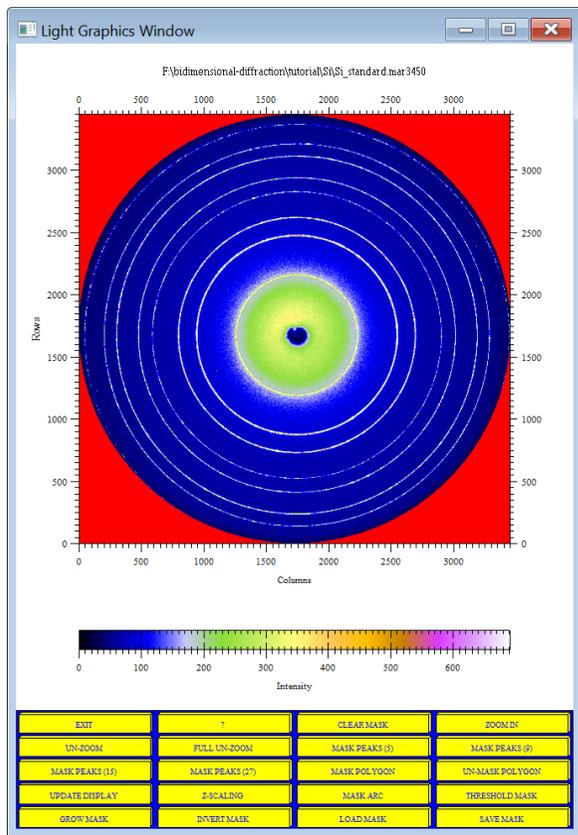
Weak peaks



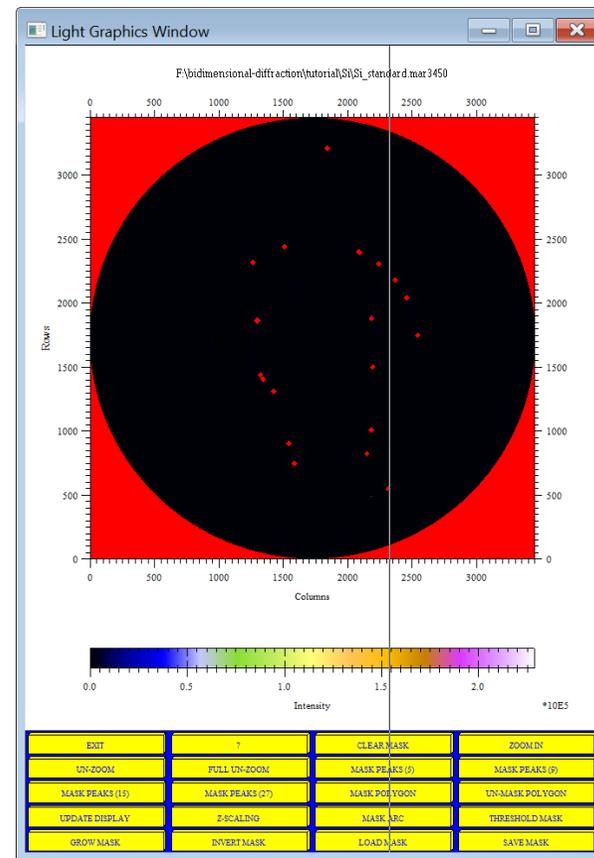
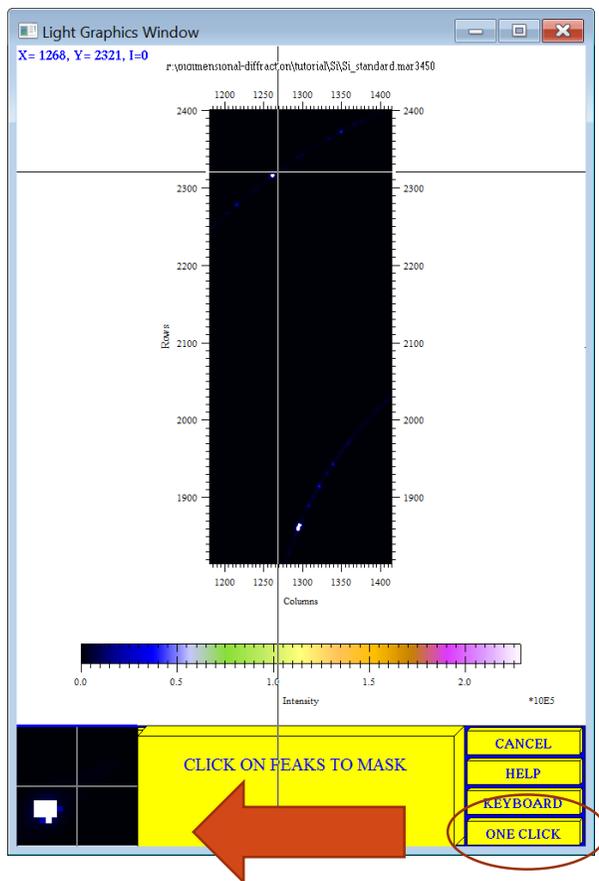
Fully automatics

Need to mask something

# Mask



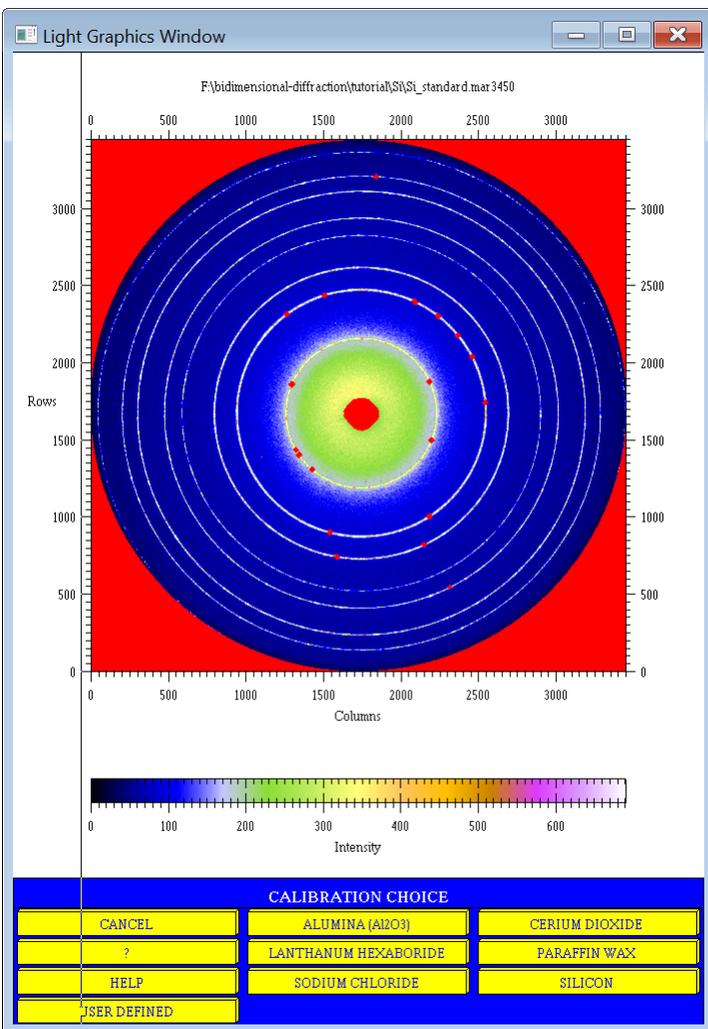
Low threshold mask



# Fully masked image

- Calibrant menu  
set of standard compound for calibration  
User could use a custom calibrant need  
a list of d-spacing in a text file

In our case Si

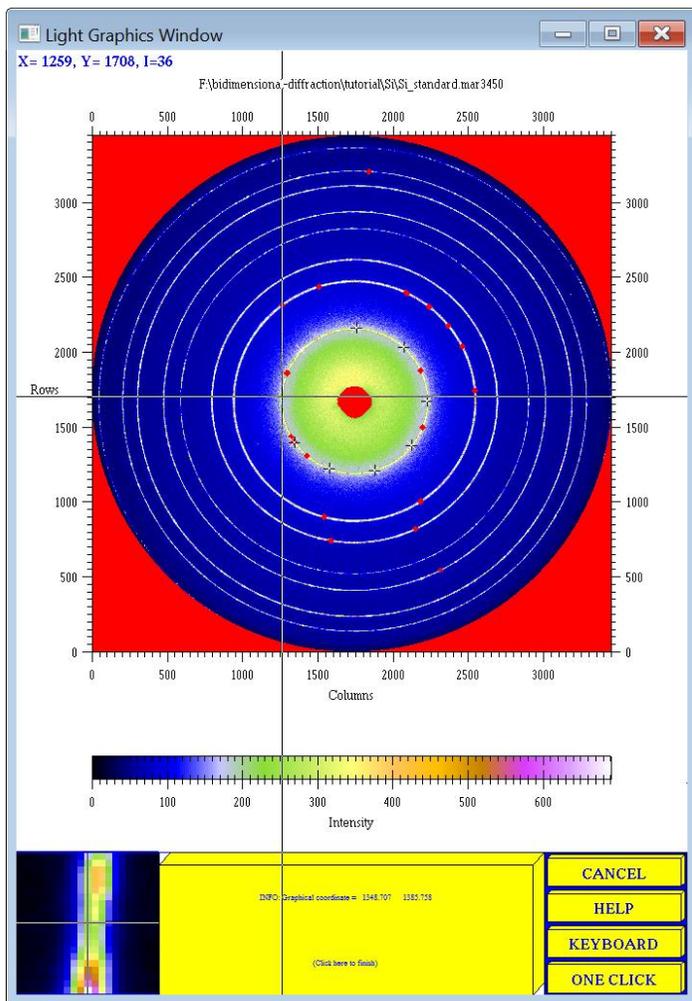


# Calibrant

O.K.	CANCEL	?	HELP	INFO
DESCRIPTIONS	VALUES	CHANGE		
SAMPLE TO DETECTOR DISTANCE (MM) (STARTING)	36.00000	DISTANCE		
WAVELENGTH (ANGSTROMS) (STARTING)	0.124619	WAVELENGTH		
SIZE OF HORIZONTAL PIXELS (MICRONS)	100.0000	X-PIXEL SIZE		
SIZE OF VERTICAL PIXELS (MICRONS)	100.0000	Y-PIXEL SIZE		
NUMBER OF AZIMUTHAL SECTIONS	360	ANGULAR SECTIONS		
REJECT OUT-LYING POSITIONS AND RE-REFINE	YES	REJECT OUTLIERS		
REJECT LIMIT FROM IDEAL (STANDARD DEVIATIONS)	4.000000	REJECT LIMIT		
OUTPUT FULL INFORMATION	YES	FULL INFO		
REFINE X/Y BEAM CENTRE	YES	REFINE BEAM X-Y		
REFINE SAMPLE TO DETECTOR DISTANCE	YES	REFINE DISTANCE		
REFINE X-RAY WAVELENGTH	NO	REFINE WAVELENGTH		
REFINE DETECTOR NON-ORTHOGONALITY	YES	REFINE TILT		
FIT INTERMEDIATE NUMBER OF RINGS	YES	EXTRA ITERATIONS		

- Approximate sample to detector distance nobody knows where is the active layer
- Wavelengths precise a possible
- Pixel size
- Number of azimuthal section number of slice used to refine the geometry stat with few (90) and repeat with 360
- Reject outline Yes but depends
- Refine beam center Yes
- Refine sample to detector distance Yes
- Refine wavelength NO correlate with distance
- Refine Tilts Yes
- Fit intermediate number of ring Depends

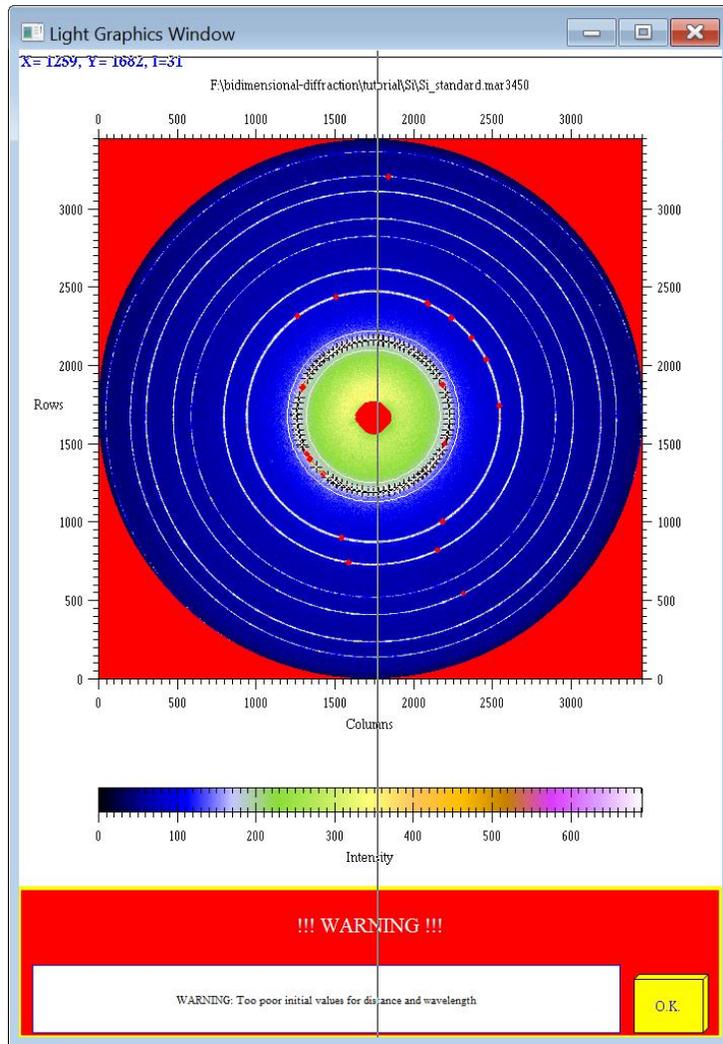
# Center of the beam



- Indicate first circle position click all around if circles not complete search to use the largest path
- Use two click
- If using user calibrant is better to write in the file the first intense reflection d-spacing

# Refinement

- Poor starting value....  
But better values are estimated  
is shell windows



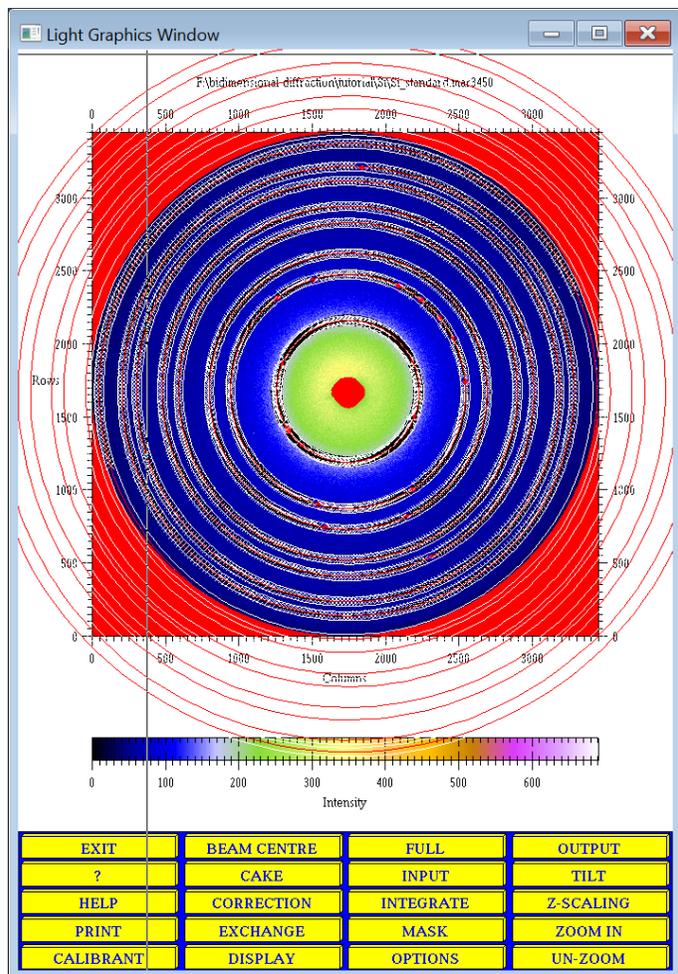
Sélectionner fit2d

```
INFO: Best fit angle of axis 1 (degrees) = -40.83157
INFO: Estimated coordinate radial position error (mm) = 0.2709064E-01
INFO: Estimated coordinate radial position error (X pixels) = 0.2709064

INFO: Calculating centre of gravity coordinates on 1 powder ring
INFO: Ring 1, Number of coordinates = 77
INFO: Number of acceptable coordinates on ring = 77
INFO: Fitting ellipse to centre of gravity coordinates
INFO: Number of coordinates = 77
INFO: Best fit ellipse centre (X/Y mm) = 174.2982 167.5236
INFO: Best fit ellipse centre (X/Y pixels) = 1742.982 1675.236
INFO: Best fit radius 1, radius 2 (mm) = 48.47878 48.47137
INFO: Best fit radius 1 (X pixels) = 484.7878
INFO: Best fit radius 2 (Y pixels) = 484.7137
INFO: Best fit angle of axis 1 (degrees) = 24.45689
INFO: Estimated coordinate radial position error (mm) = 0.2733386E-01
INFO: Estimated coordinate radial position error (X pixels) = 0.2733386
INFO: Estimated sample to detector distance = 362.978 mm
WARNING: The calculated sample to detector distance (given the initial
estimate of the wavelength) is too different from the user
estimate of the wavelength. Please check the initial values
of both the distance and the wavelength and alter them using
the graphical entry form. Also check that you are using the
correct ring.
```

# calibration

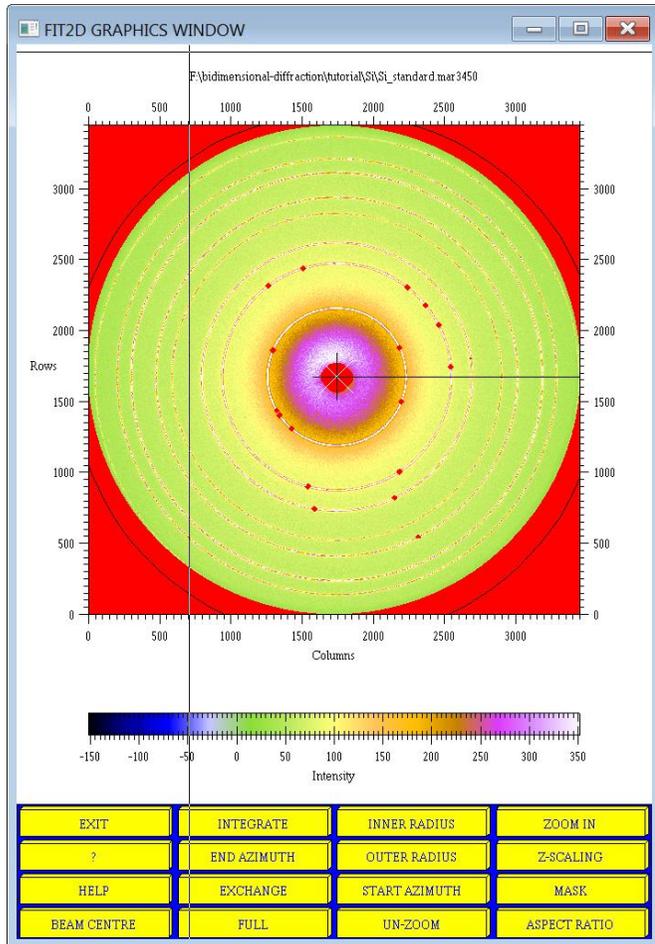
Need to check, cake it



Selectionner fit2d

```
INFO: Ring 6, Number of coordinates = 360
INFO: Ring 7, Number of coordinates = 357
INFO: Ring 8, Number of coordinates = 175
INFO: Ring 9, Number of coordinates = 0
INFO: Ring 10, Number of coordinates = 0
INFO: Ring 11, Number of coordinates = 0
INFO: Ring 12, Number of coordinates = 0
INFO: Ring 13, Number of coordinates = 0
INFO: Ring 14, Number of coordinates = 0
INFO: Ring 15, Number of coordinates = 0
INFO: Number of function calls = 7
INFO: Sum of squares = 5.0928
INFO: Number of rejected coordinates = 0
INFO: Number of function calls = 7
INFO: Sum of squares = 5.0928
INFO: Refined Beam centre = 1742.917 1675.293 (pixels)
INFO: Refined Beam centre = 174.292 167.529 (mm)
INFO: Refined sample to detector distance = 363.401 mm
INFO: Refined wavelength = 0.41594 Angstroms
INFO: Energy (keU) = 29.80844
INFO: Refined tilt plane rotation angle = -166.370 degrees
INFO: Refined tilt angle = -0.054 degrees
INFO: ROT X = 0.052 ROT Y = 0.013 degrees
INFO: Stability indicator (proportional to D-spacing; Angstroms) = 1.55360
```

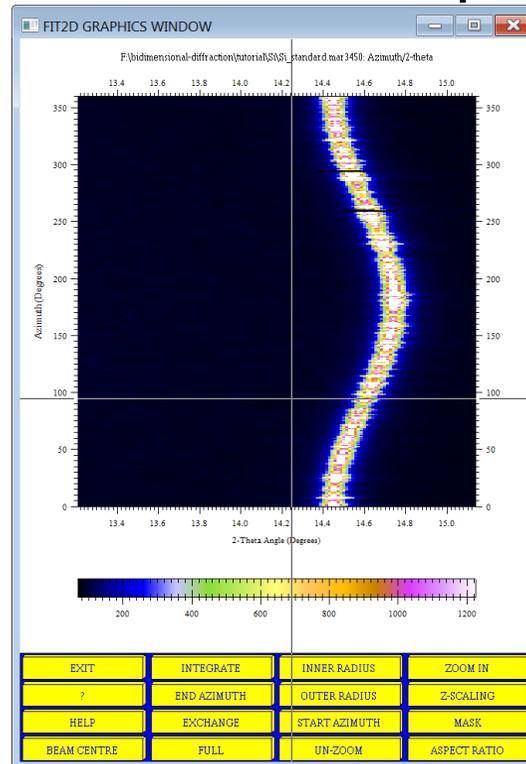
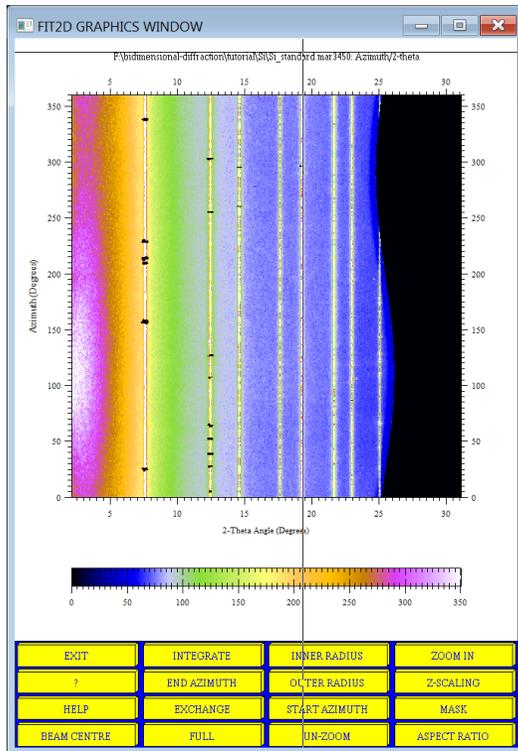
# Cake



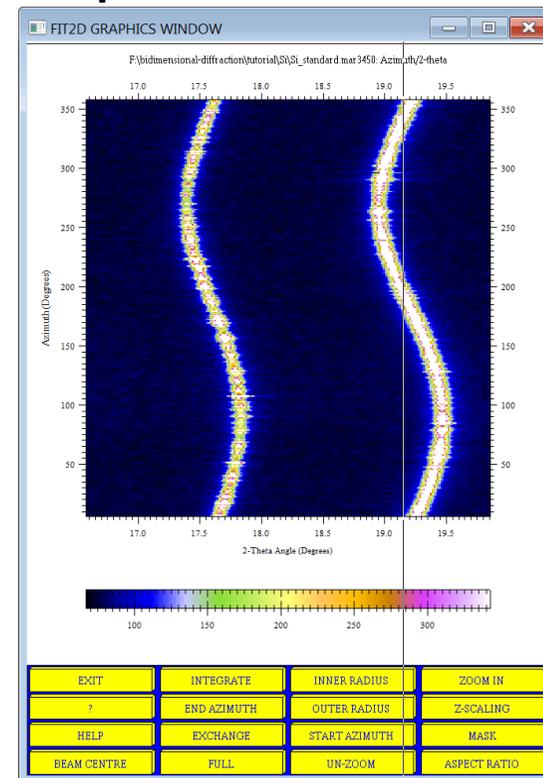
- We define an azimuthal start ,stop
- Inner radius outer radius
- Number of azimuthal bins
- Number of 2theta bins
- Intensity conservation
- Polarization correction
- polarization factor
- Geometrical correction for intensity

# Azimuthal plot

- Azimuthal plot represent line → OK

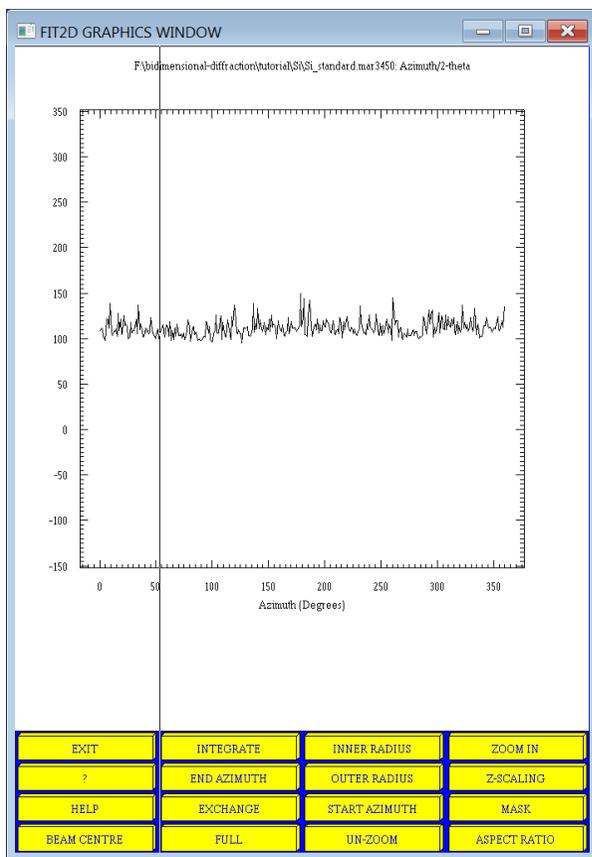


Wrong center

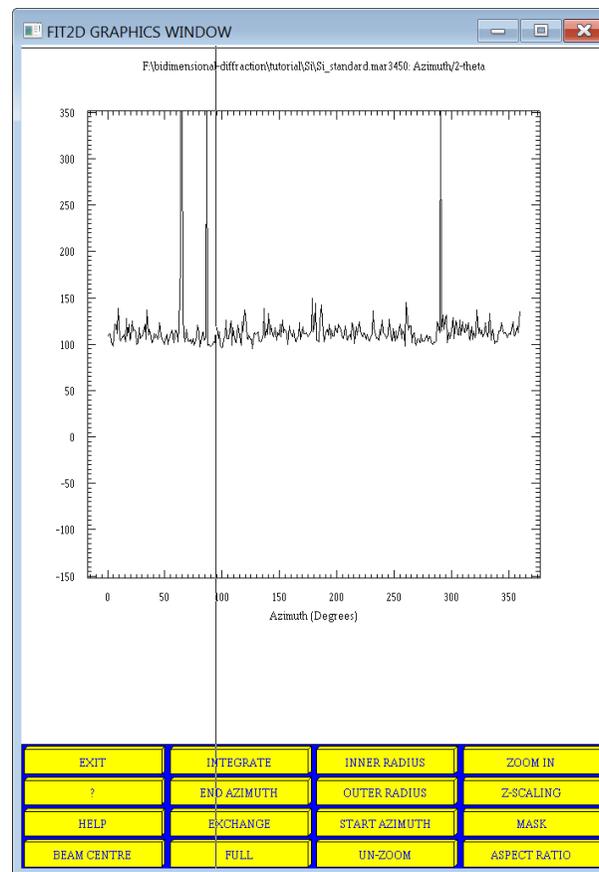


Wrong rot or tilt angle

# Integral of azimuthal scan



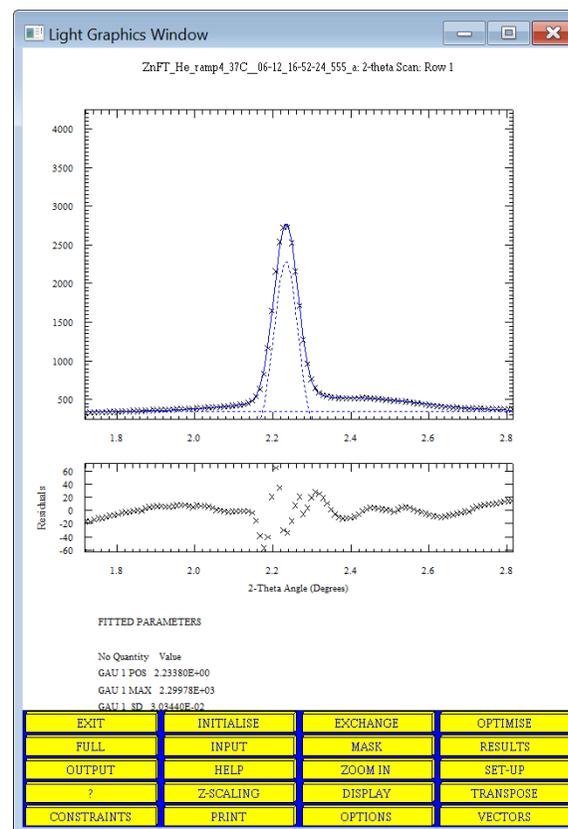
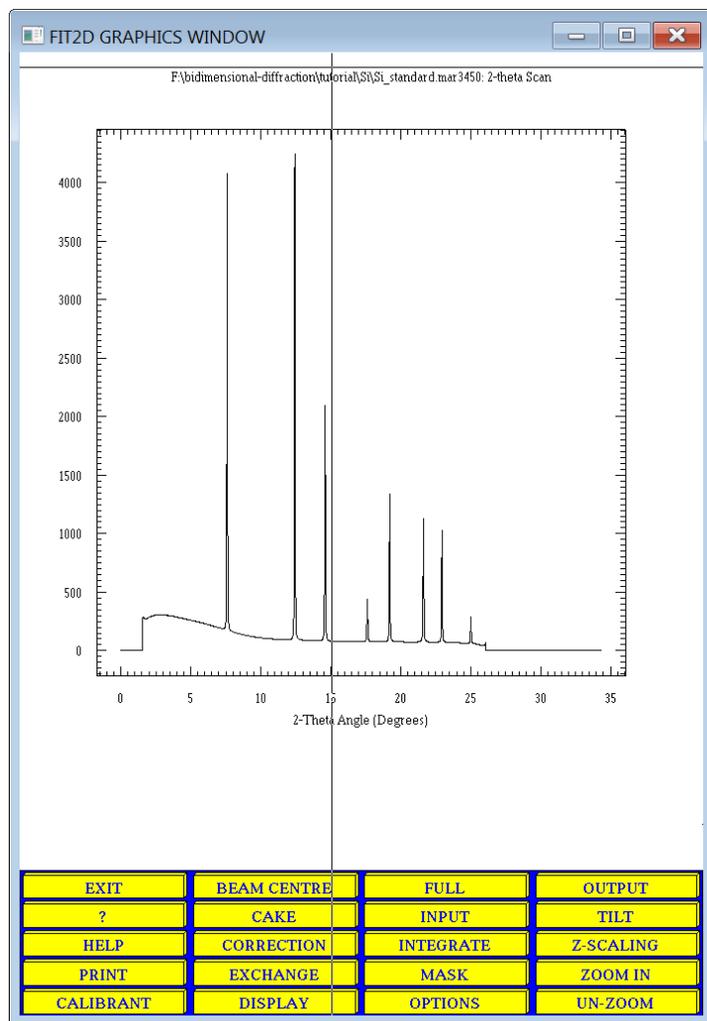
Good mask



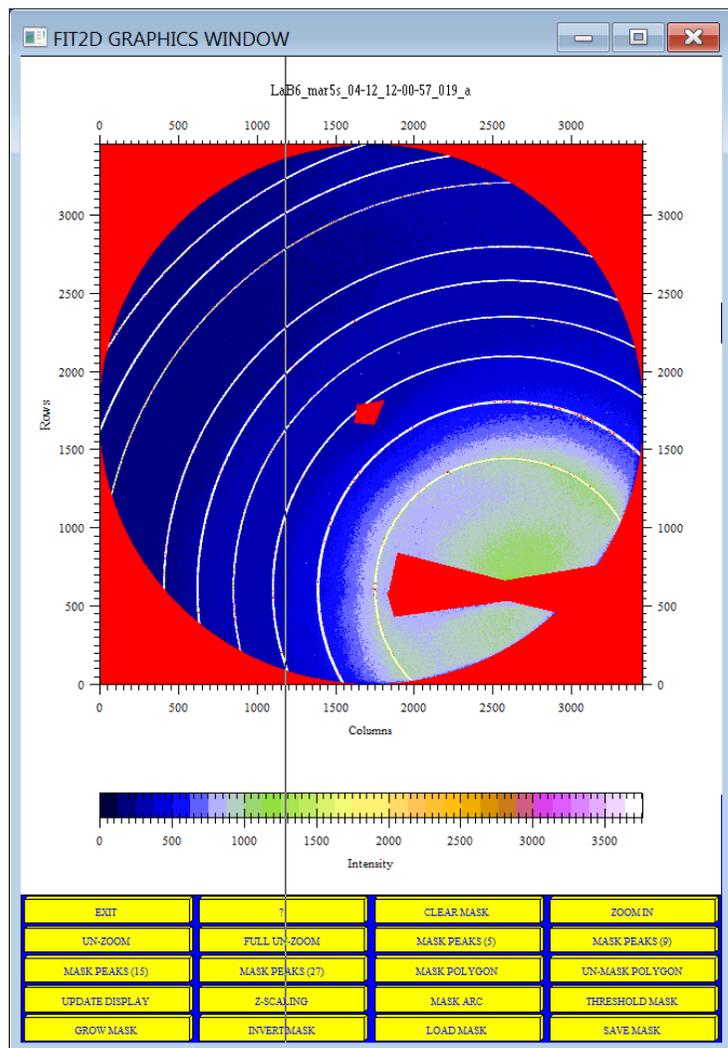
Few outliers

# Integration

- Possibility of save or fit peaks also iteratively with mfit



# Macros



start with calibration

Exit

Macros/logfile Create a macro Exit

Powder Diffraction

Input (Open a file of the batch)

Integrate, output, exit

Macros/logfile STOP macro Exit

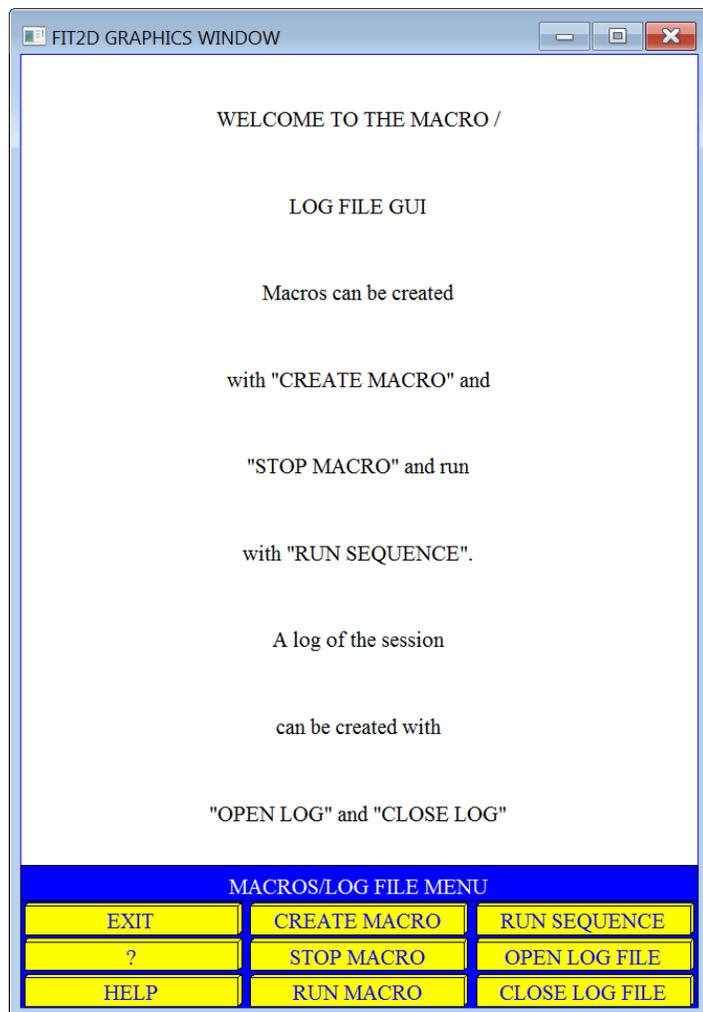
# Macro files

```
%!*\ This is a comment line  
EXIT  
ON-LINE CRYSTALLOGRAPHY  
INPUT  
F:\ZnFT_72_2\ZnFT_He_r_555_a.edf  
EXIT  
POWDER DIFFRACTION (2-D)  
INTEGRATE  
O.K.  
O.K.  
OUTPUT  
CHIPLOT  
O.K.  
EXCHANGE  
EXIT  
MACROS / LOG FILE  
%!*\ END OF IO MACRO FILE
```

Change  
input file>>>#IN  
Output>>>#OUT

```
%!*\ This is a comment line  
EXIT  
ON-LINE CRYSTALLOGRAPHY  
INPUT  
#IN  
EXIT  
POWDER DIFFRACTION (2-D)  
INTEGRATE  
O.K.  
O.K.  
OUTPUT  
CHIPLOT  
O.K.  
EXCHANGE  
EXIT  
MACROS / LOG FILE  
%!*\ END OF IO MACRO FILE
```

# Run sequence



Run sequence

First file of the batch, last file, increment

And wait

# Alternative option for non sequential name

- Write a file with all the instruction or generate it with sed and awk

%!\ BEGINNING OF GUI MACRO FILE	O.K.	O.K.
%!\	O.K.	OUTPUT
%!\ This is a comment line	OUTPUT	CHIPLLOT
%!\	CHIPLLOT	O.K.
EXIT	O.K.	INPUT
POWDER DIFFRACTION (2-D)	INPUT	F:\Data-SOLEIL-CRISTAL-20130419-
INPUT	F:\Data-SOLEIL-CRISTAL-20130419-	dec13\Audebrand-
F:\Data-SOLEIL-CRISTAL-20130419-	dec13\Audebrand-	MAR345\ZnFT_72_2\ZnFT_He_ramp4_104C_
dec13\Audebrand-	MAR345\ZnFT_72_2\ZnFT_He_ramp4_100C_	_06-12_18-18-59_590_a.edf
MAR345\ZnFT_72_2\ZnFT_But_RT_06-12_19-	_06-12_18-14-12_588_a.edf	O.K.
43-12_619_a.edf	O.K.	INTEGRATE
O.K.	INTEGRATE	O.K.
INTEGRATE	O.K.	O.K.
O.K.	O.K.	OUTPUT
O.K.	OUTPUT	CHIPLLOT
OUTPUT	CHIPLLOT	O.K.
CHIPLLOT	O.K.	INPUT
O.K.	INPUT	F:\Data-SOLEIL-CRISTAL-20130419-
INPUT	F:\Data-SOLEIL-CRISTAL-20130419-	dec13\Audebrand-
F:\Data-SOLEIL-CRISTAL-20130419-	dec13\Audebrand-	MAR345\ZnFT_72_2\ZnFT_He_ramp4_108C_
dec13\Audebrand-	MAR345\ZnFT_72_2\ZnFT_He_ramp4_100C_	_06-12_18-21-19_591_a.edf
MAR345\ZnFT_72_2\ZnFT_He_ramp4_100C_	_06-12_18-16-38_589_a.edf	O.K.
_06-12_18-11-53_587_a.edf	O.K.	INTEGRATE
O.K.	INTEGRATE	O.K.
INTEGRATE	O.K.	

# Further reading

- Powder Diffraction Theory and Practice  
Chapter 14 Two-dimensional Diffraction Using Area Detectors  
HINRICHSEN, DINNEBIER AND JANSEN  
(first part of presentation is mainly based on it)
- Two-dimensional\_X-ray\_Diffraction B. B. He
- Norby, P. (1997) Synchrotron Powder Diffraction using Imaging Plates: Crystal Structure Determination and Rietveld Refinement. J App Crys 30, 21-30.
- A P Hammersley, S O Svensson, and A Thompson, H Graafsma, °A Kvick, and J P Moy, "Calibration and correction of distortions in 2D detector systems", Rev. Sci. Instr., (SRI-94), 66, 2729-2733 (1995)
- A P Hammersley, S O Svensson, M Hanfland, A N Fitch, and D Hausermann, "Two-Dimensional Detector Software: From Real Detector to Idealised Image or Two-ThetaScan", High Pressure Research, 14, 235-248 1996